

The Effect of Unresolved Contaminant Stars on the Cross- Matching of Photometric Catalogues

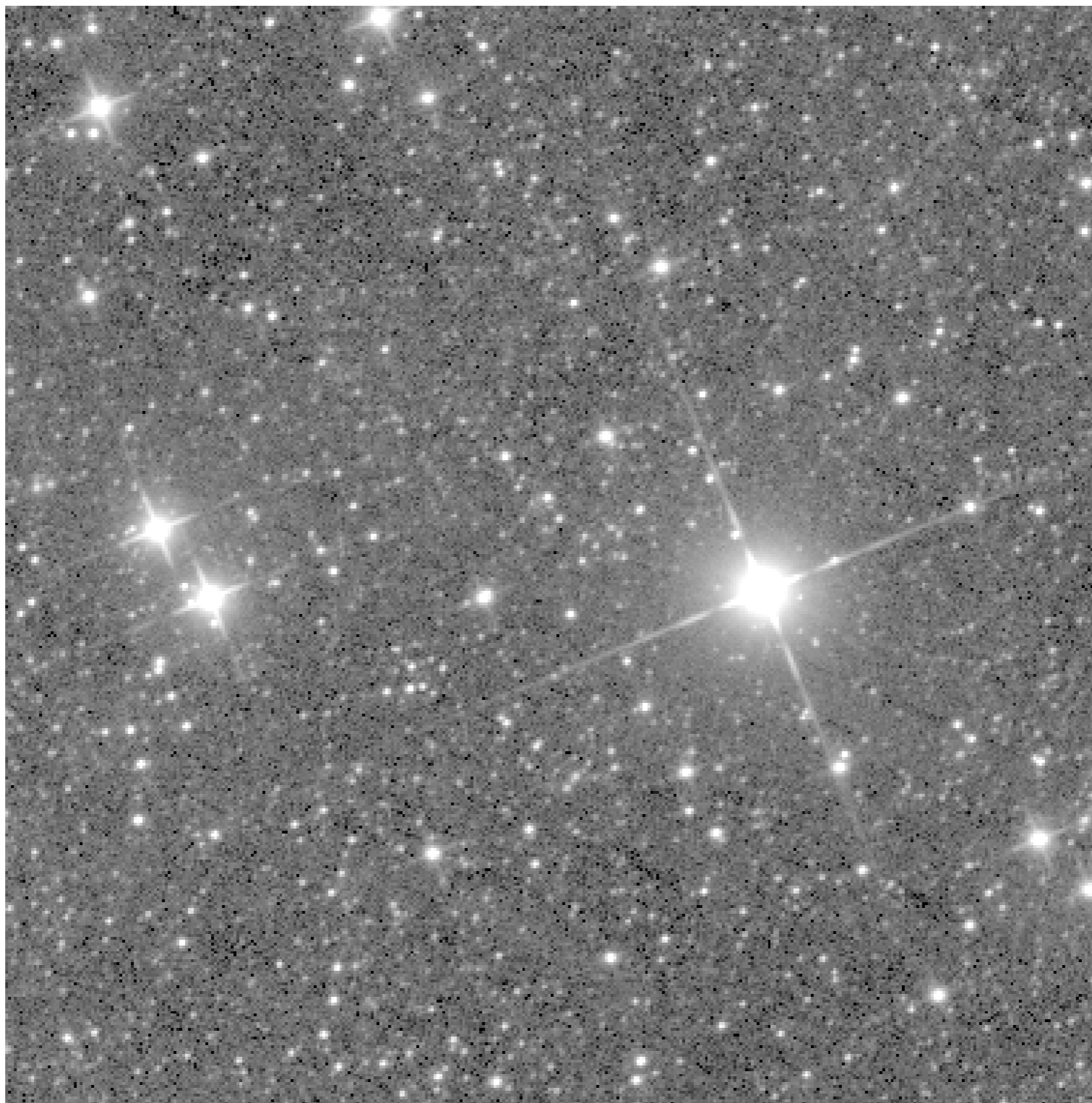
Tom Wilson, STScI

Tim Naylor, University of Exeter
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TESS Data Workshop, 2/12/19

Tom Wilson @onoastrmer

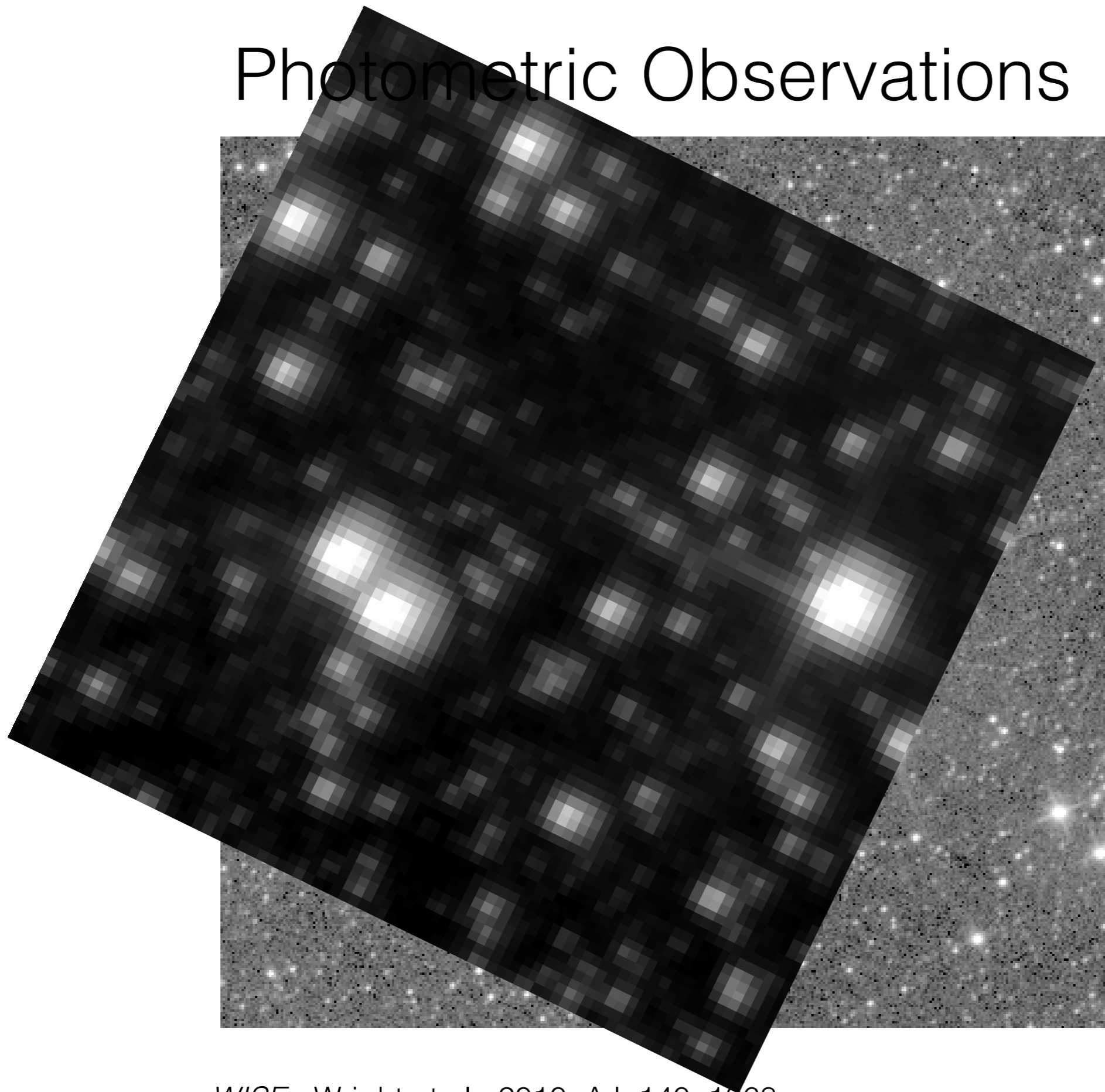
Photometric Observations



WISE - Wright et al., 2010, *AJ*, 140, 1868

WISE W1
Tom Wilson @onoastrmer

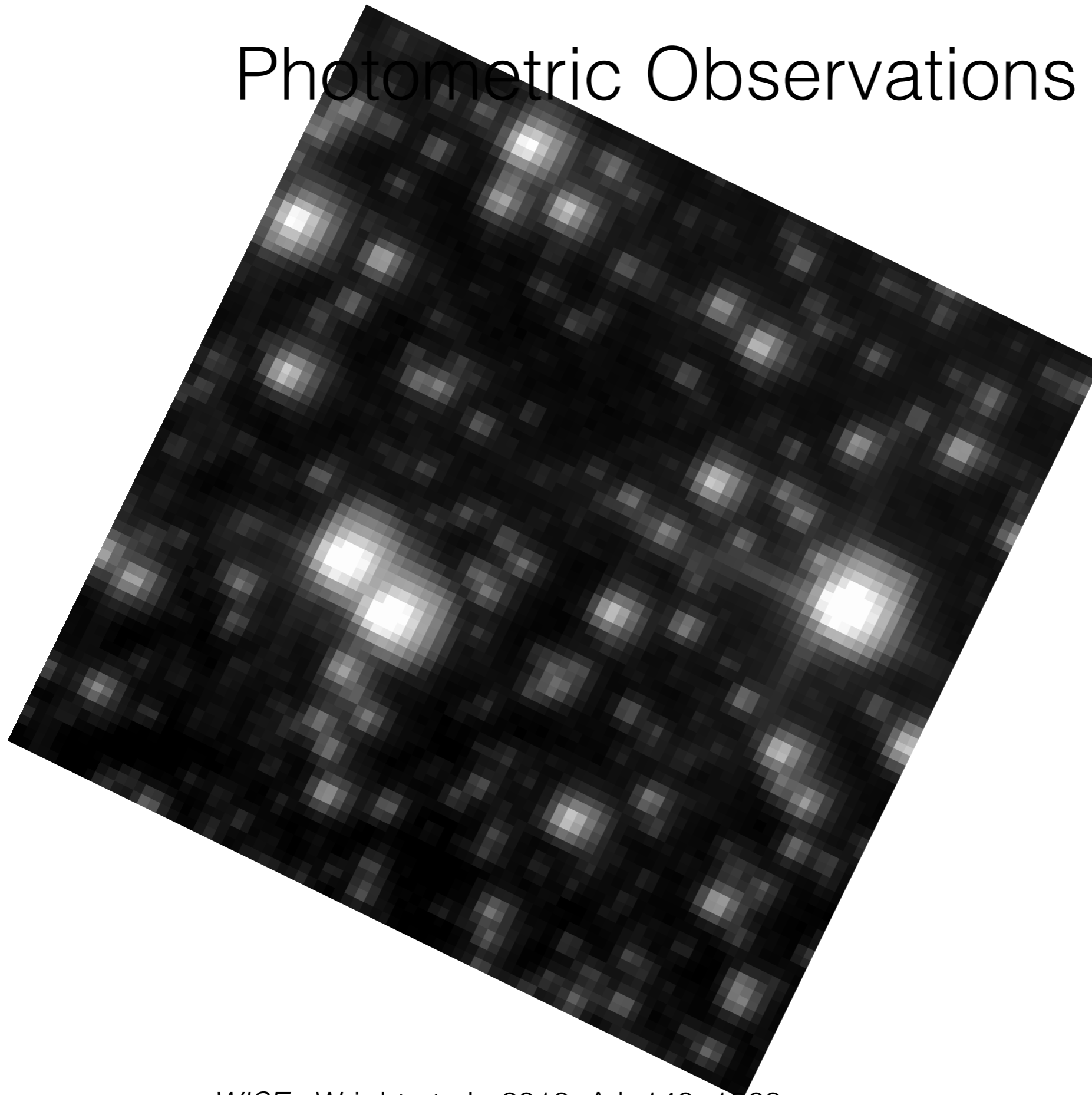
Photometric Observations



WISE - Wright et al., 2010, *AJ*, 140, 1868
TESS - Ricker et al., 2015, *JATIS*, 1, 14003

TESS T
Tom Wilson @onoastrmer

Photometric Observations



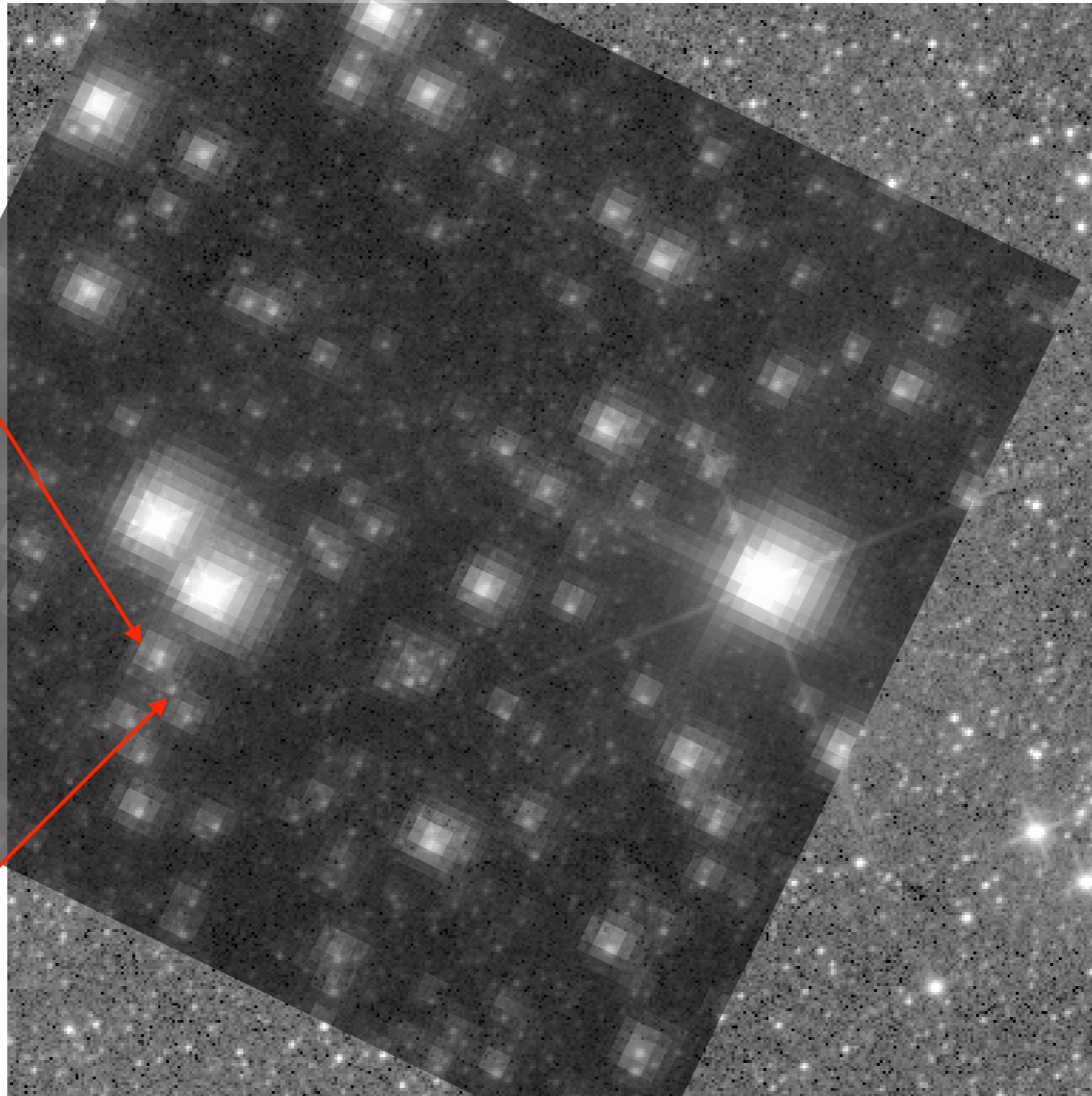
WISE - Wright et al., 2010, *AJ*, 140, 1868
TESS - Ricker et al., 2015, *JATIS*, 1, 14003

TESS T
Tom Wilson @onoastrmer

Photometric Observations

Unresolved
double star
“pair”!

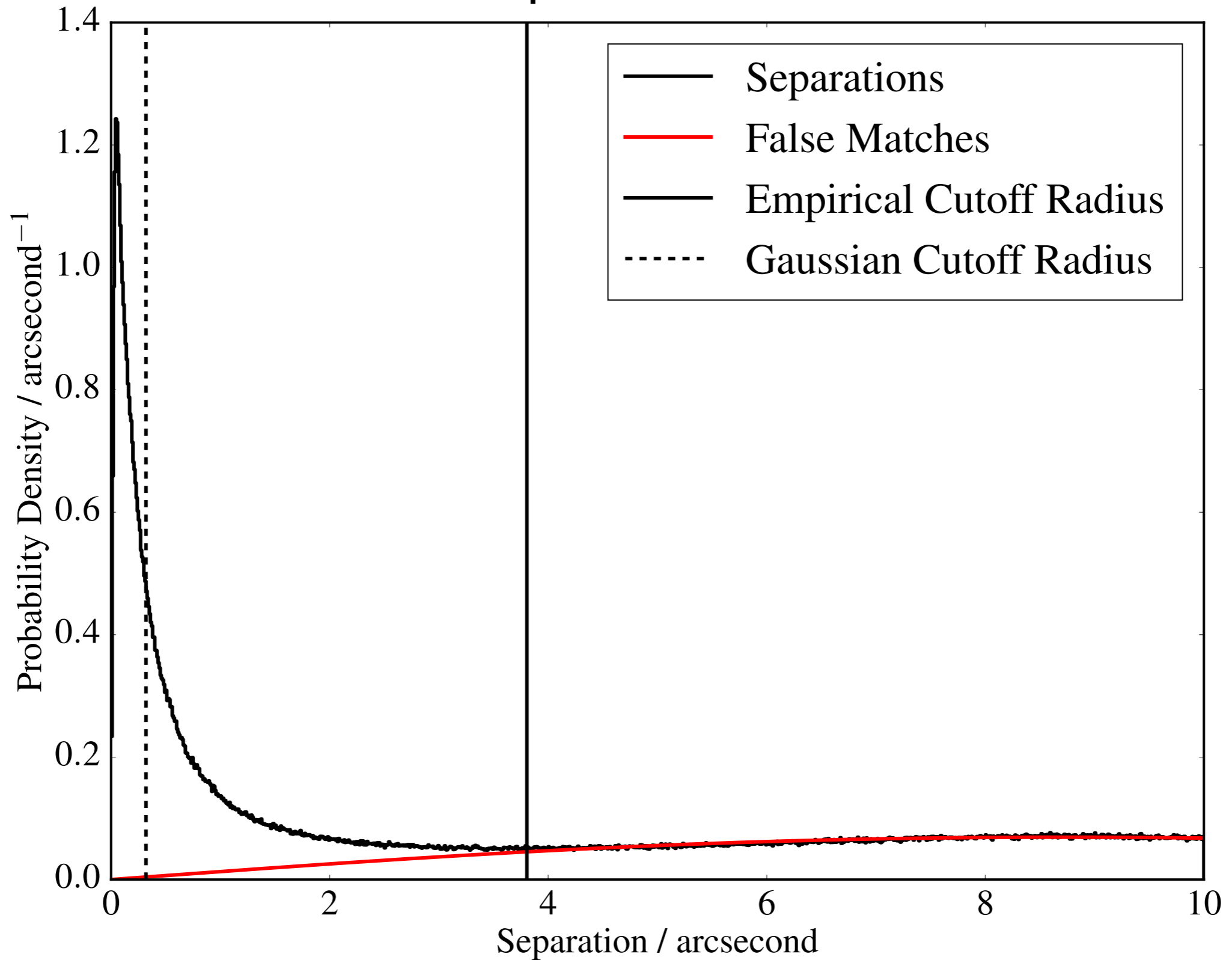
Completely
unresolved
source!



WISE - Wright et al., 2010, *AJ*, 140, 1868
TESS - Ricker et al., 2015, *JATIS*, 1, 14003

TESS T
Tom Wilson @onoastrmer

Cross-match Separation Distributions

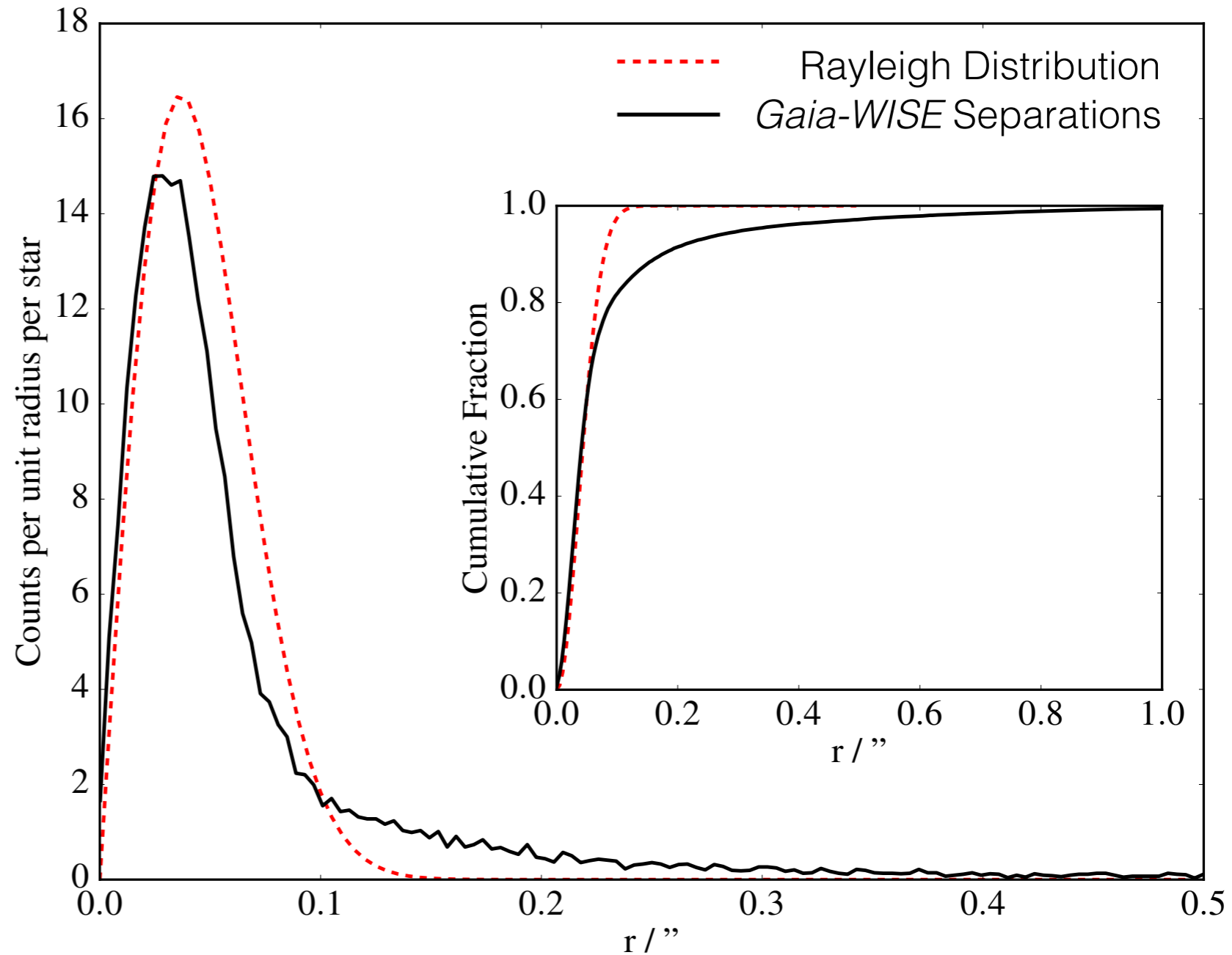
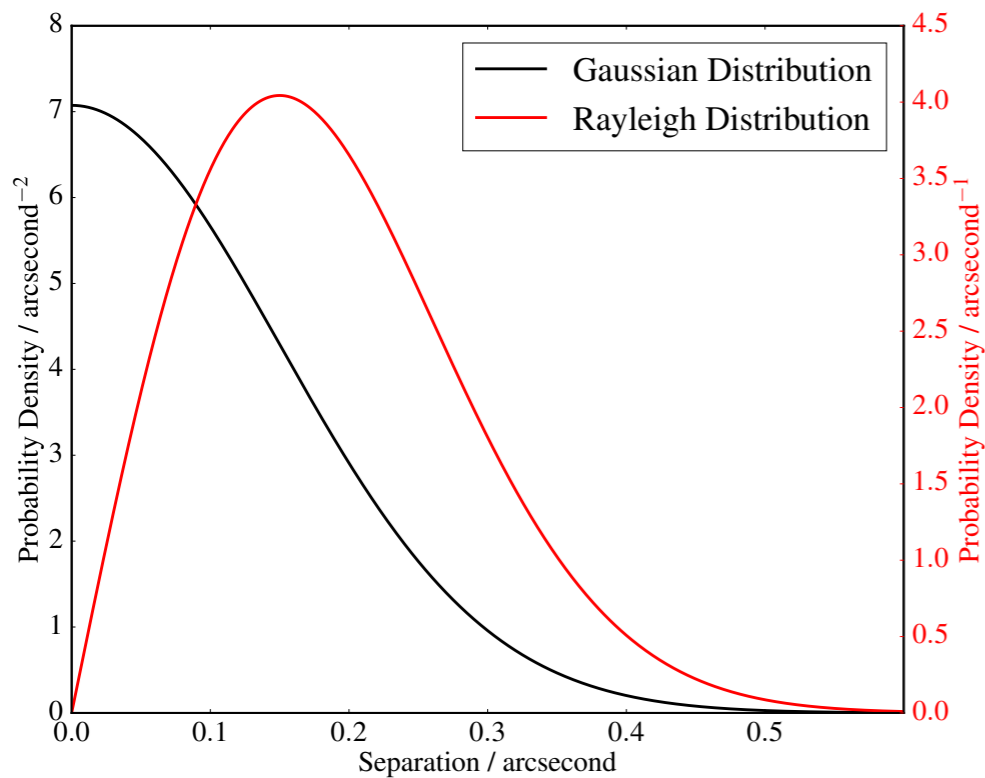


The Astrometric Uncertainty Function

$$g(x, y, \sigma) = \frac{1}{2\pi\sigma^2} e^{-\frac{x^2+y^2}{2\sigma^2}}$$

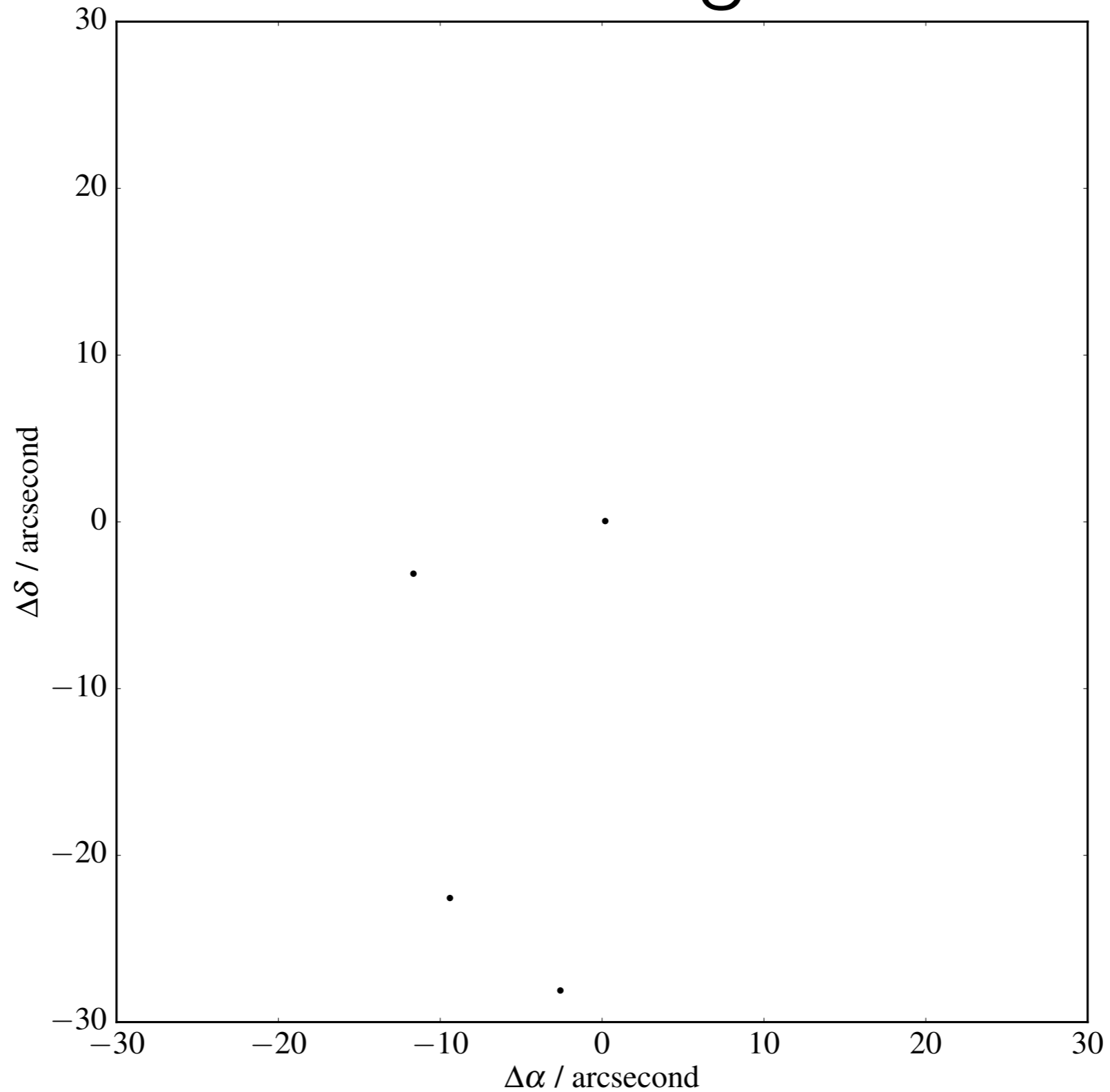
↓

$$g(r, \sigma) = \frac{r}{\sigma^2} e^{-\frac{r^2}{2\sigma^2}}$$

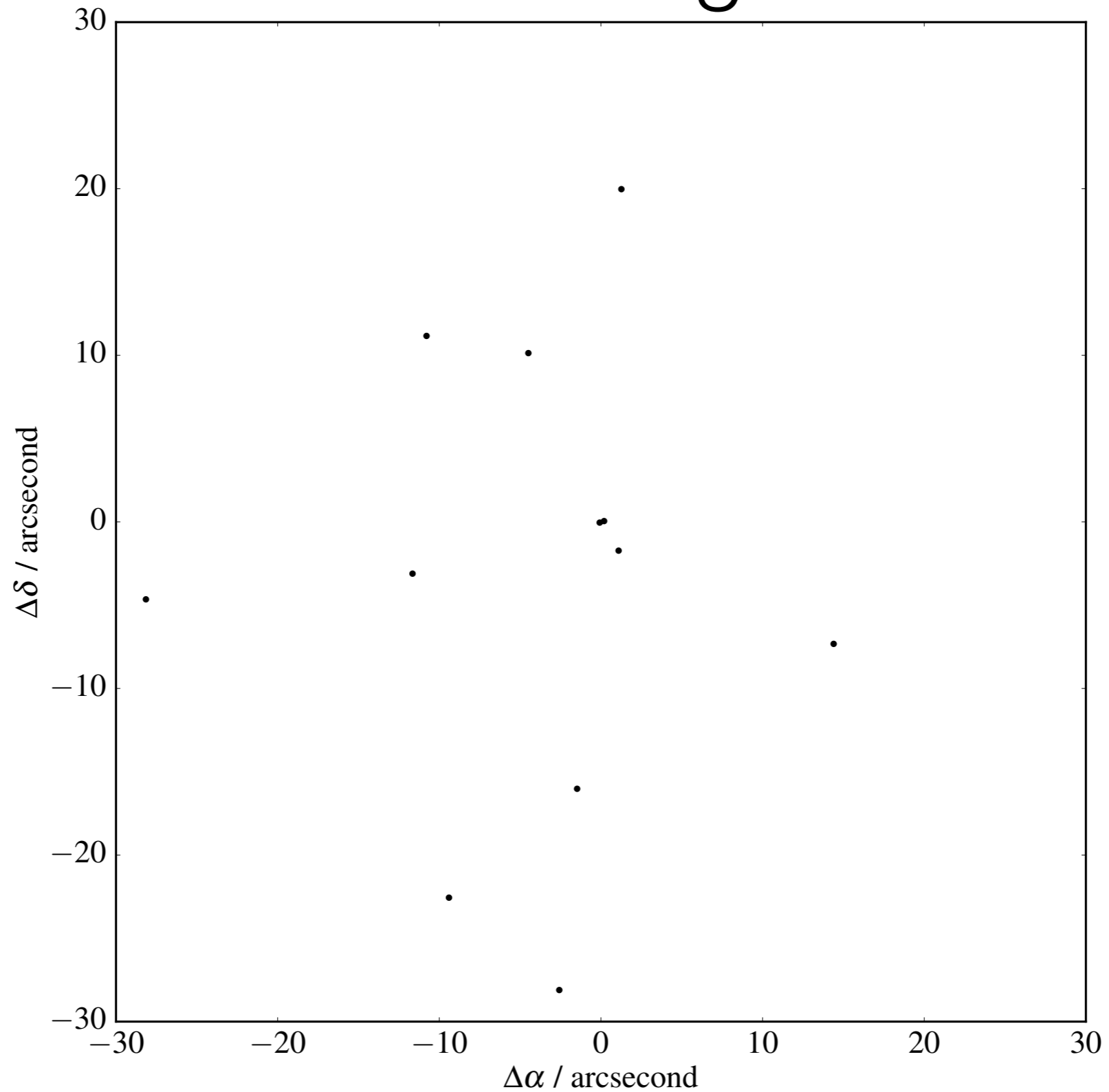


Gaia DR2 - Gaia Collaboration, Brown A. G. A., et al. 2018, A&A, 616, 1
 Wilson & Naylor, 2017, MNRAS, 468, 2517
WISE - Wright et al., 2010, AJ, 140, 1868

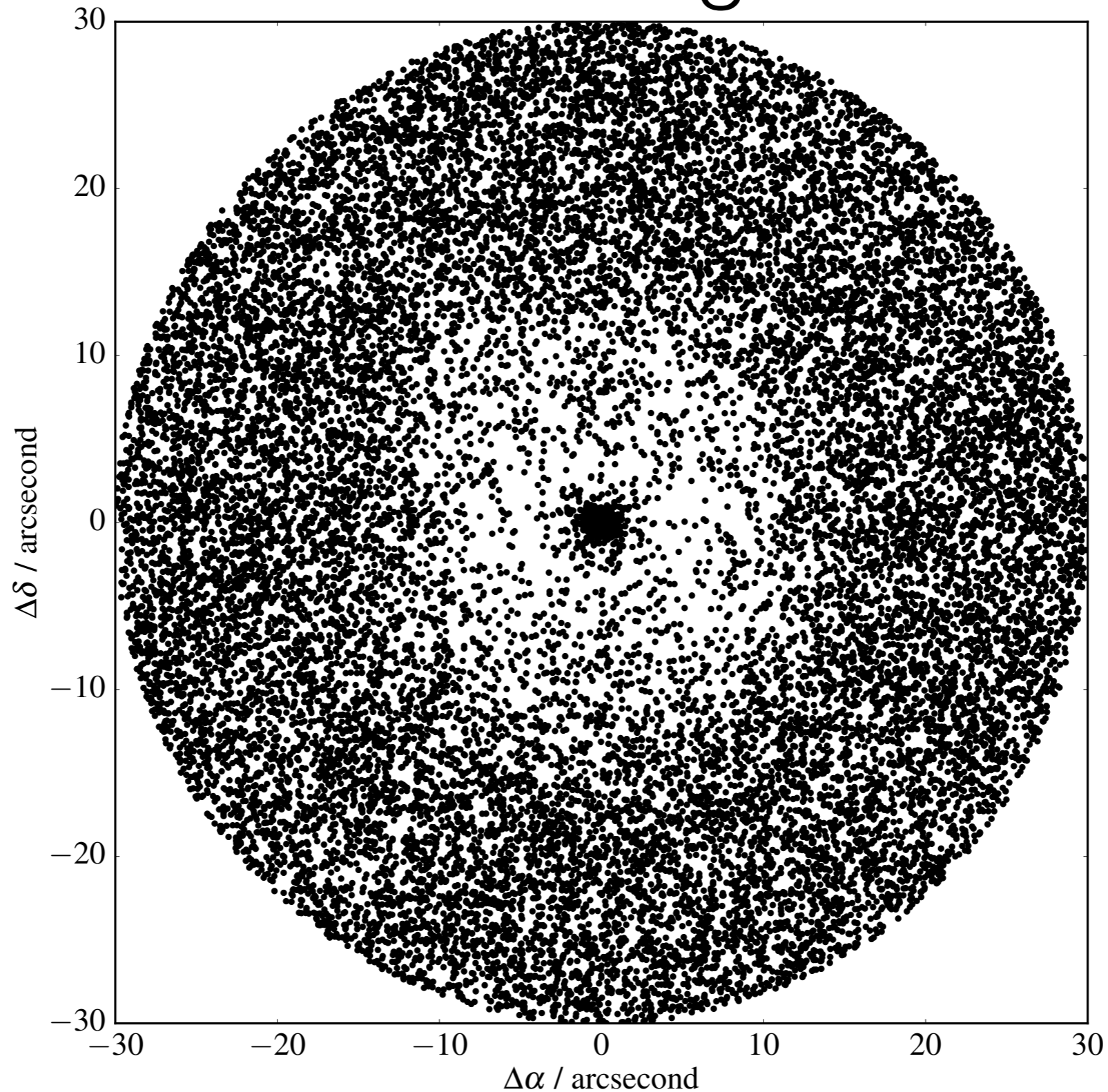
The Astrometric Uncertainty Function: Crowding



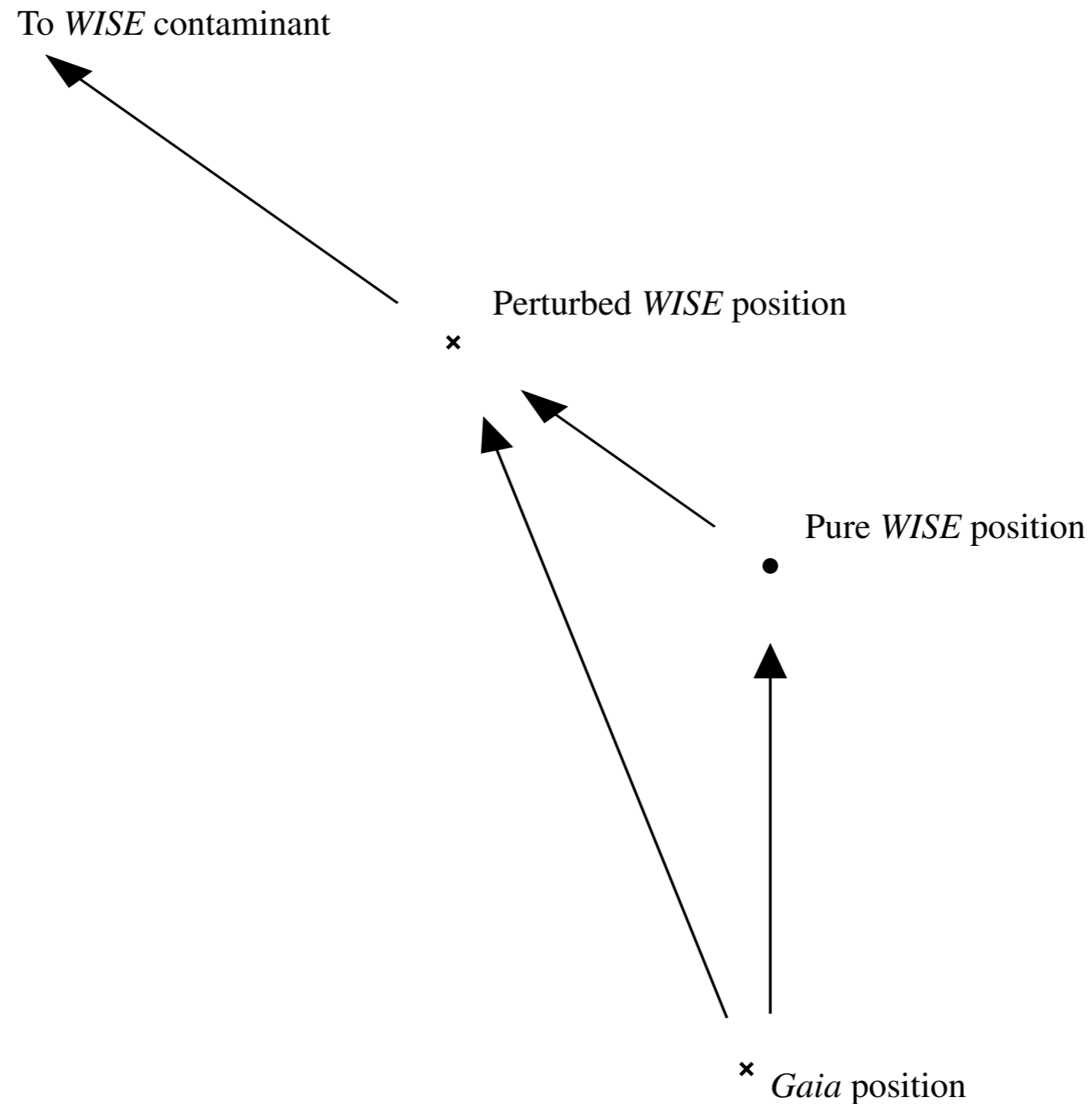
The Astrometric Uncertainty Function: Crowding



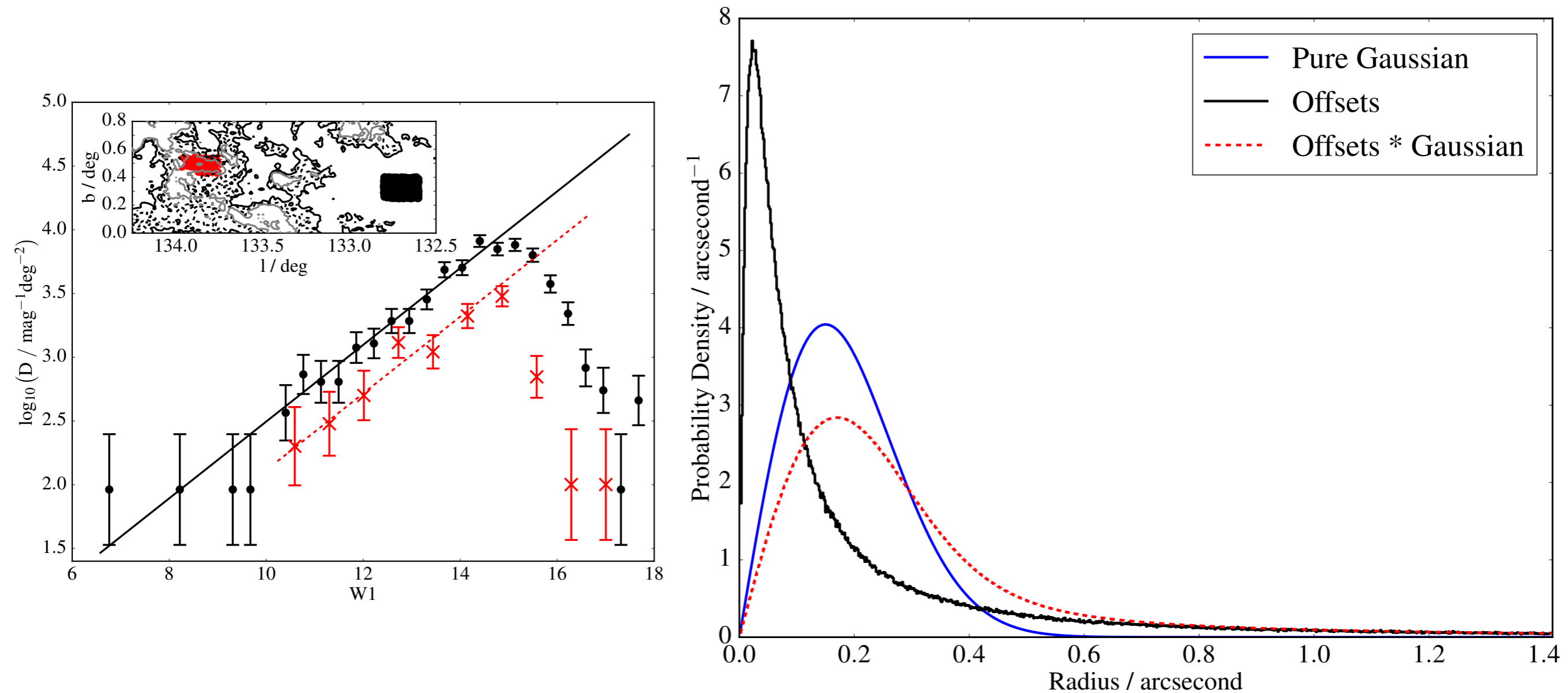
The Astrometric Uncertainty Function: Crowding



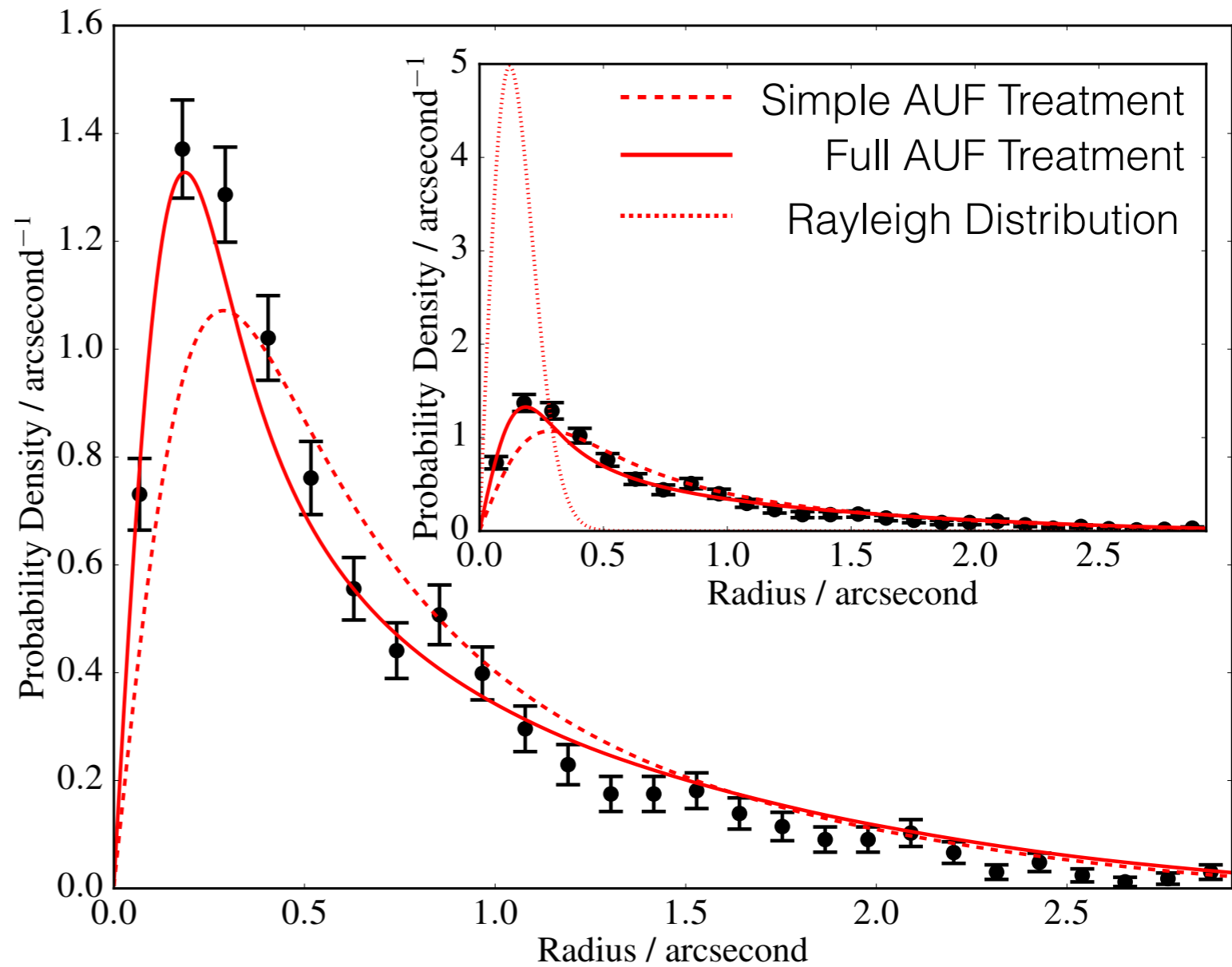
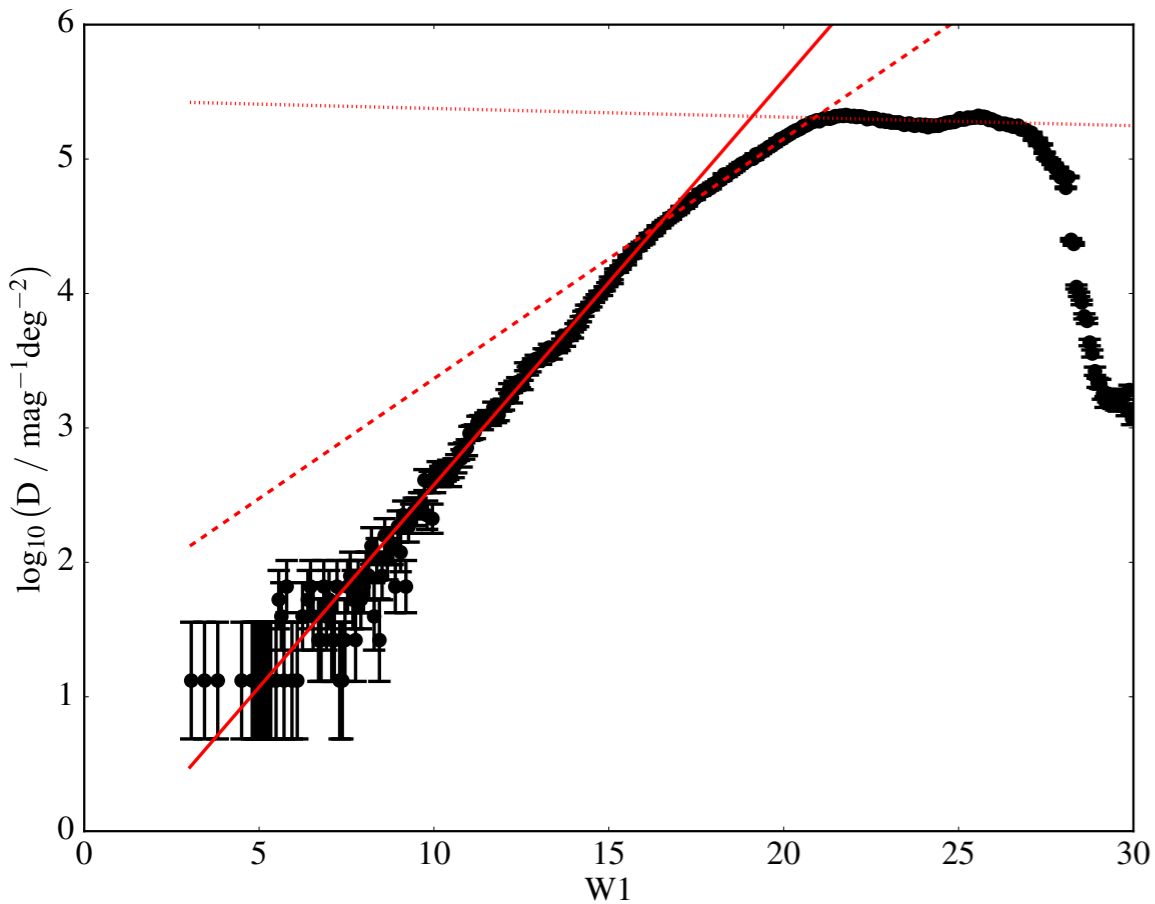
The Astrometric Uncertainty Function: Perturbation



The Astrometric Uncertainty Function: Building Empirical AUFs

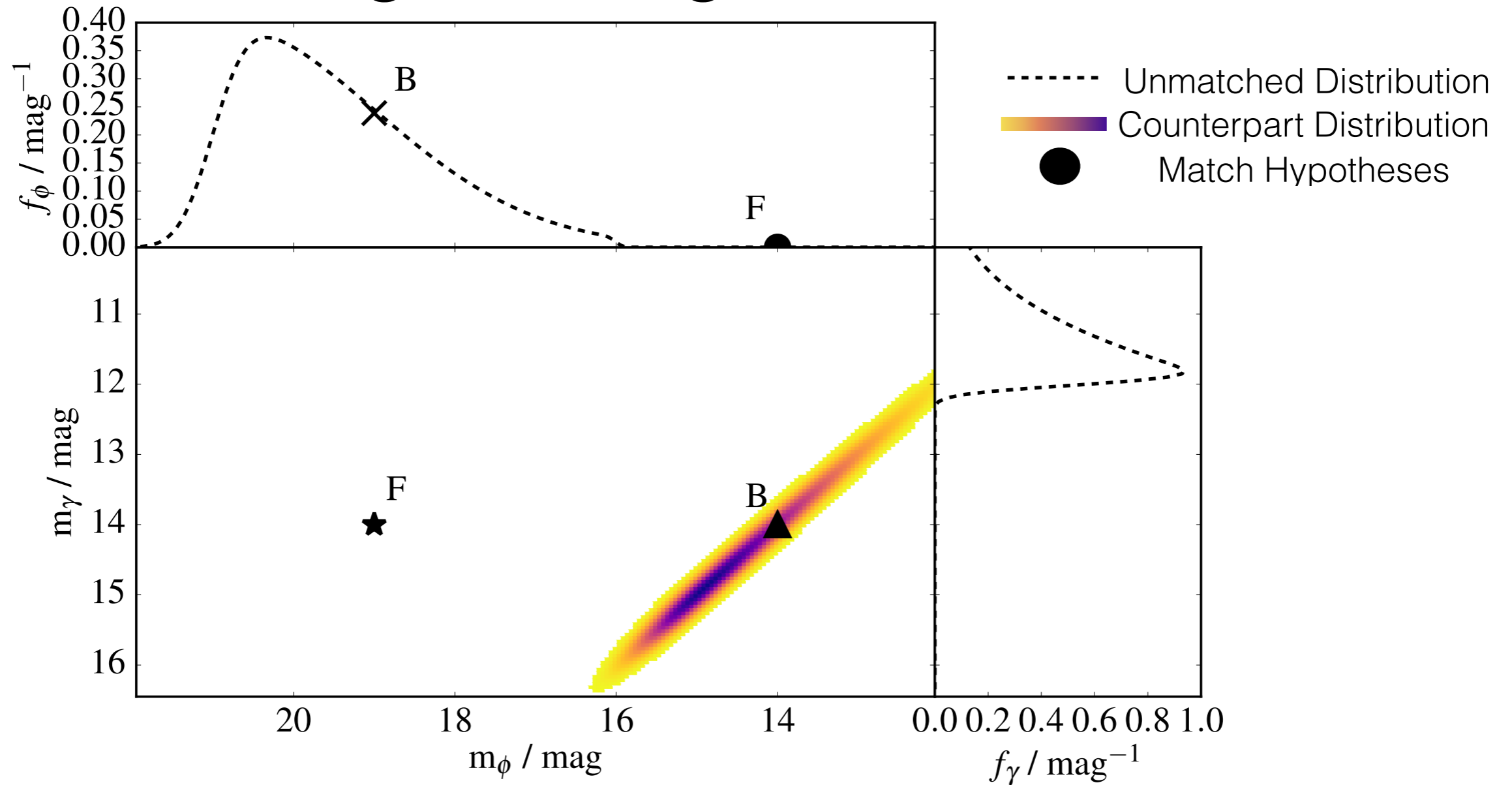


Contamination Effects: Effects Below Sensitivity Limit



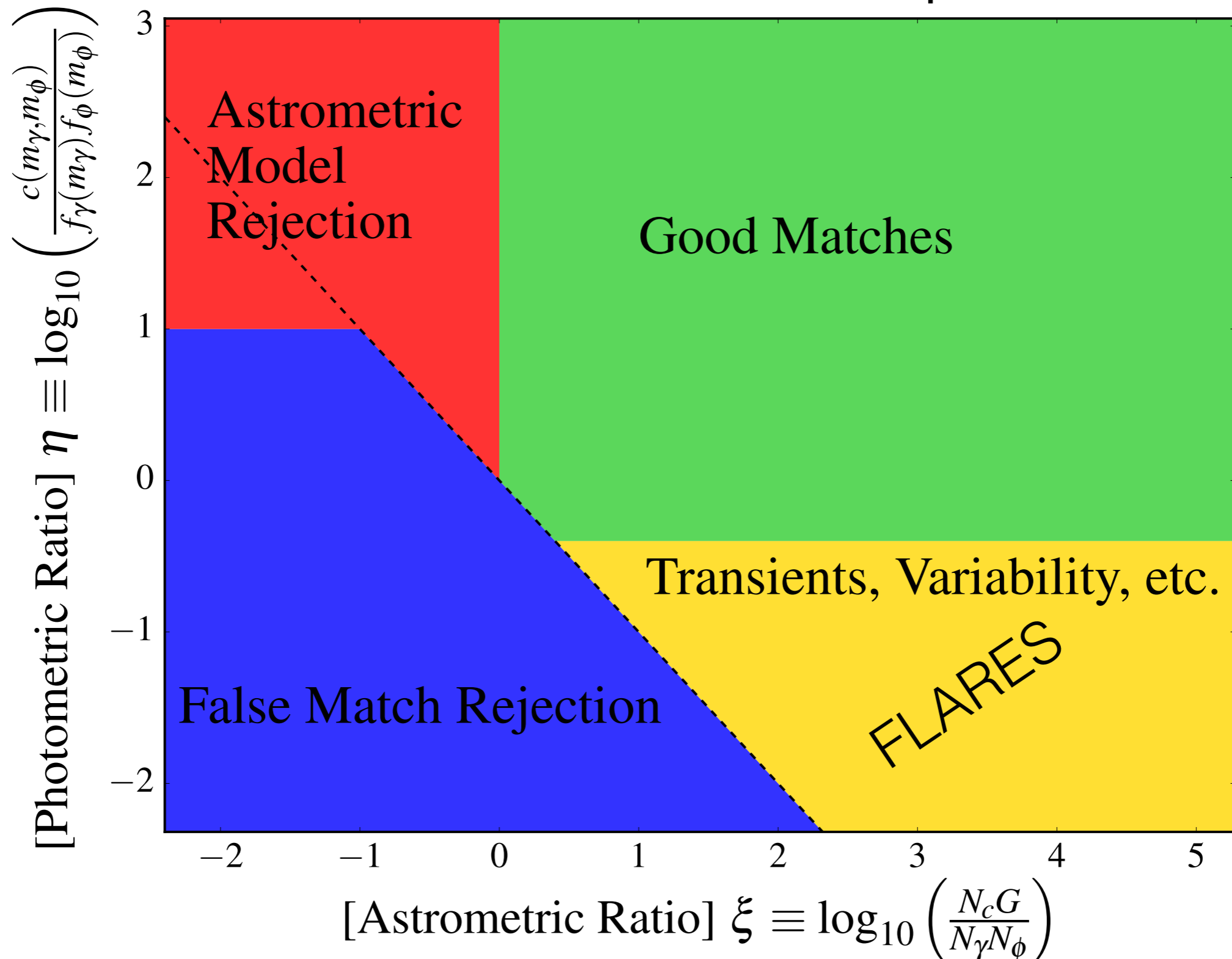
Effect crucial for
faint M dwarfs

Probability-based Catalogue Matching: Including the Magnitude Information

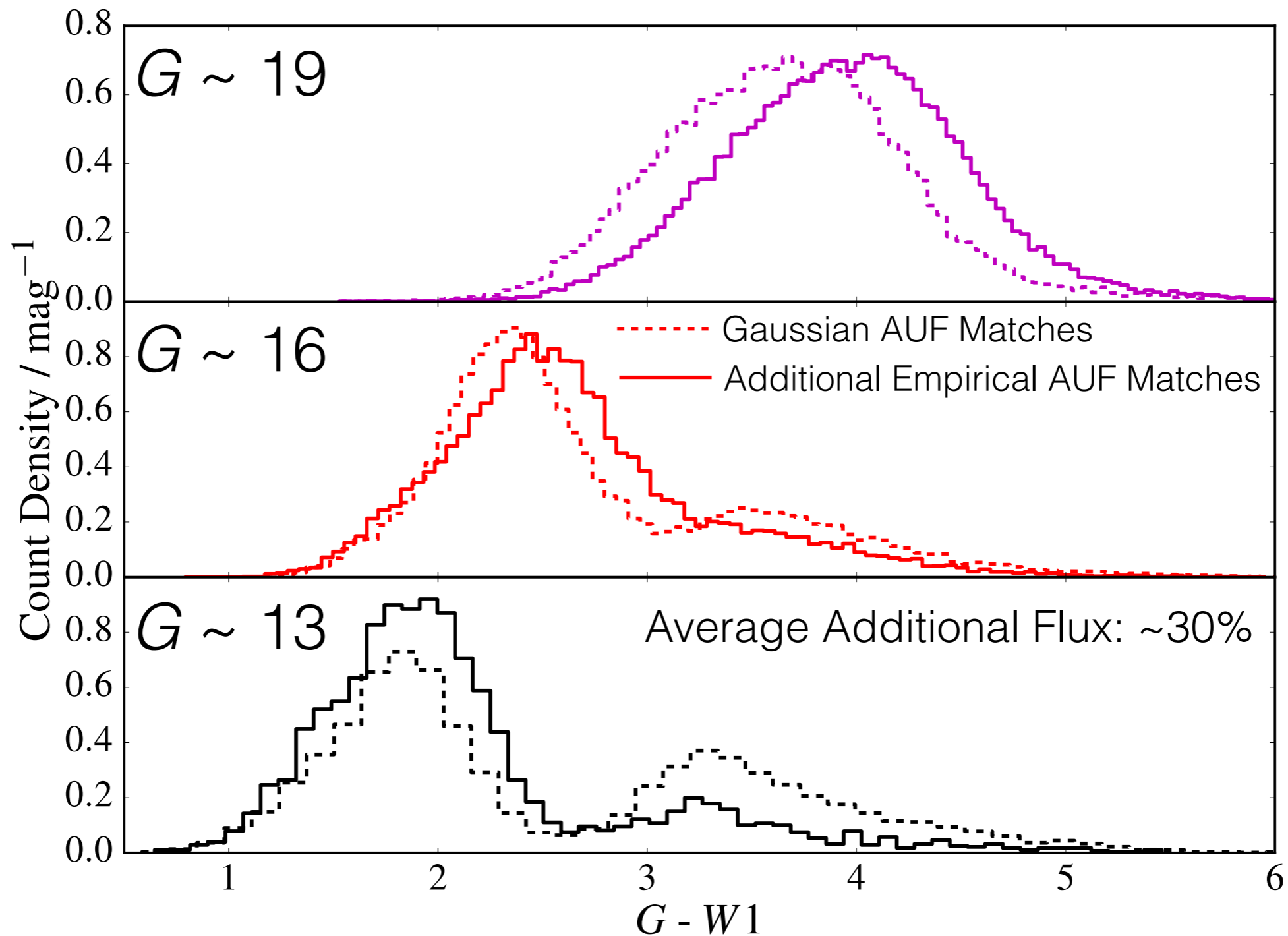


$$\begin{aligned}
 g(x_k, y_k, x_l, y_l) &= N_c \iint_{-\infty}^{+\infty} h_\gamma(\Delta x_{kl} - x, \Delta y_{kl} - y) h_\phi(x, y) dx dy \\
 &= N_c \times (h_\gamma * h_\phi)(\Delta x_{kl}, \Delta y_{kl}).
 \end{aligned}$$

Probability-based Catalogue Matching: The Likelihood Ratio Space

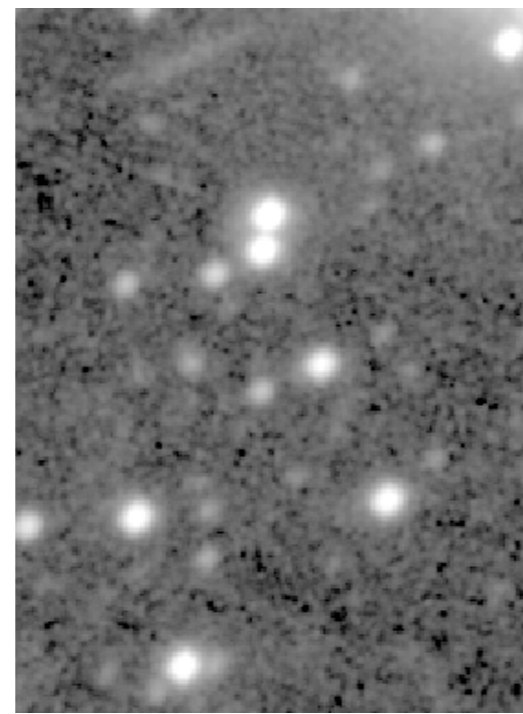
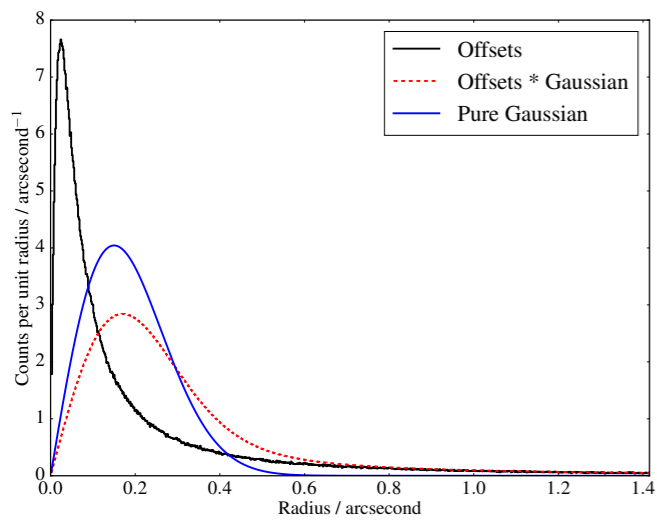


Contamination Effects: Perturbation-Colour Correlation

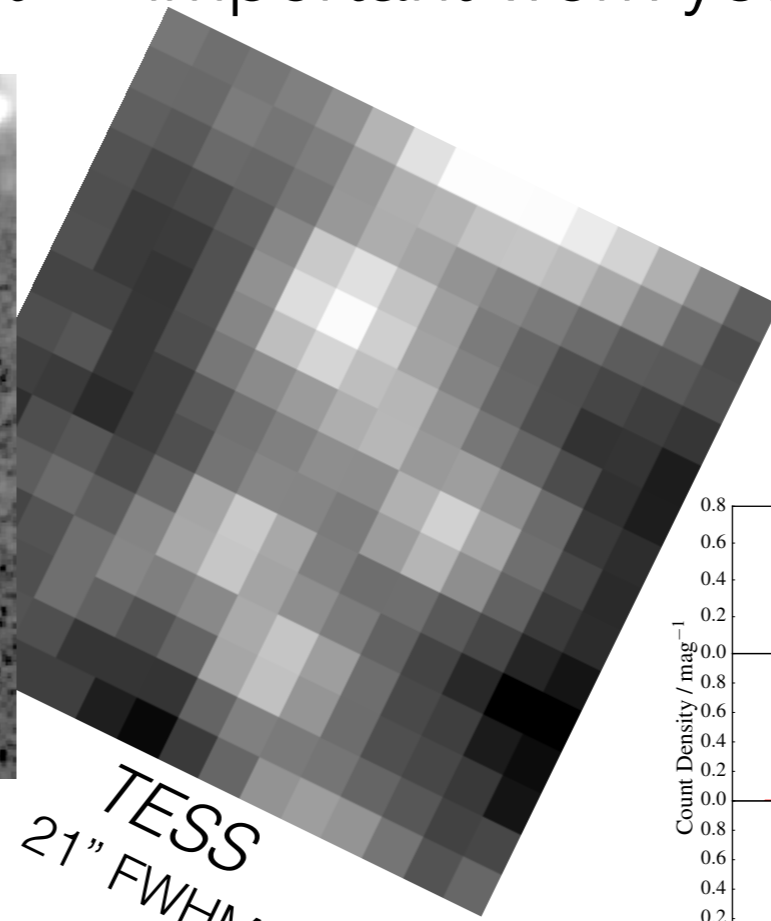


The Effects of Unresolved Contaminant Stars on the Cross-Matching of Photometric Catalogues: Conclusions

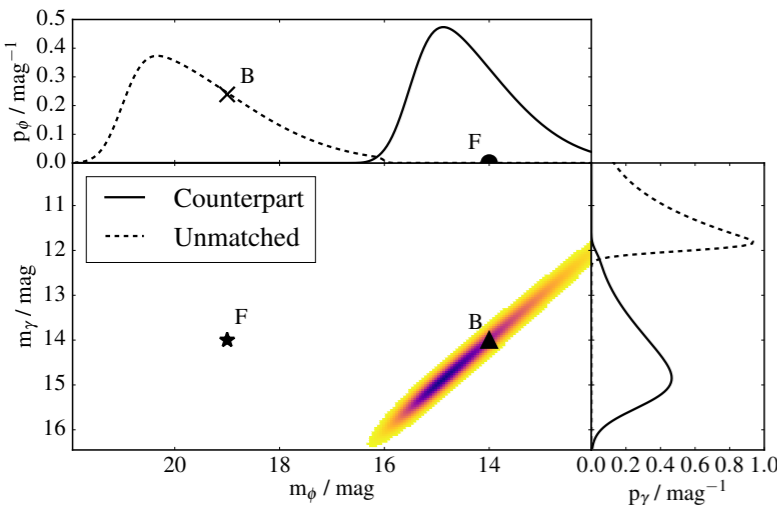
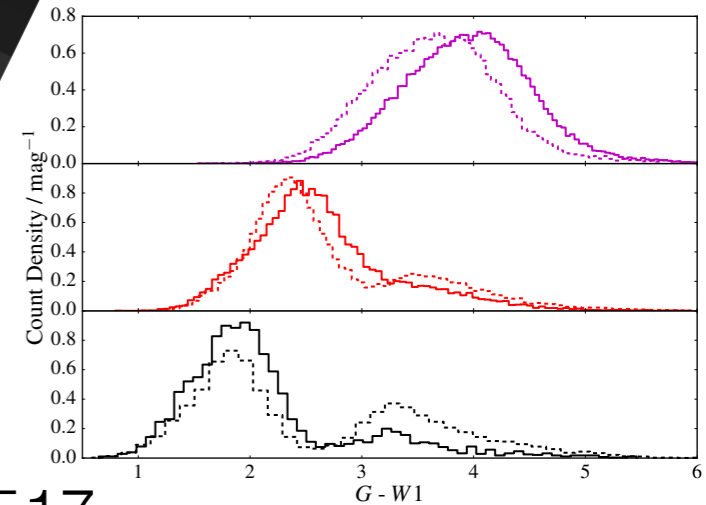
- Blended star contamination causes positional shifts
- *WISE* objects are up to 30% flux contaminated, with *TESS* suffering blending on a larger scale
- Disentangle this information with proper treatment in the cross-match to a higher angular resolution dataset — important work yet to be done!



WISE
6" FWHM



TESS
21" FWHM

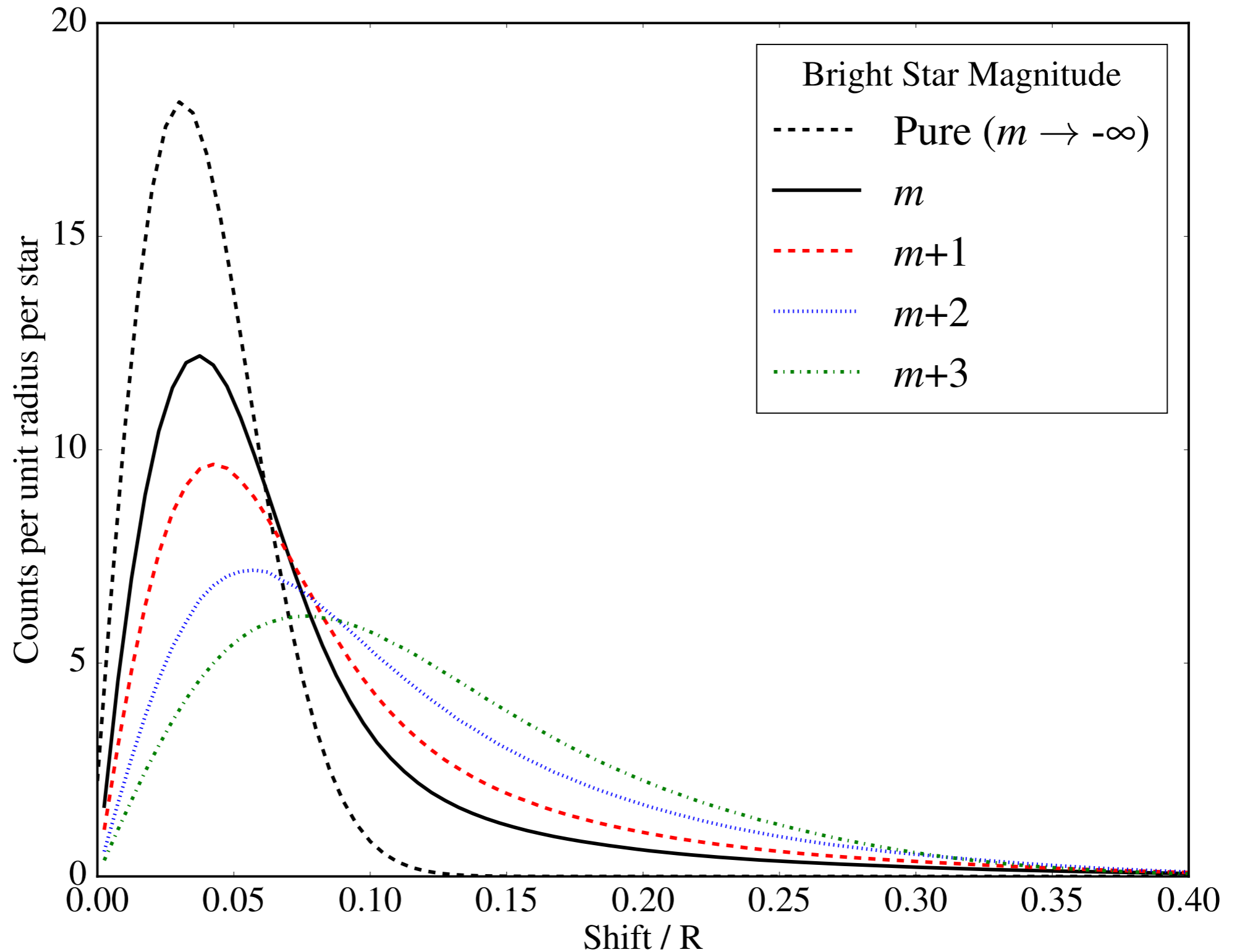


Wilson & Naylor, 2017, MNRAS, 468, 2517

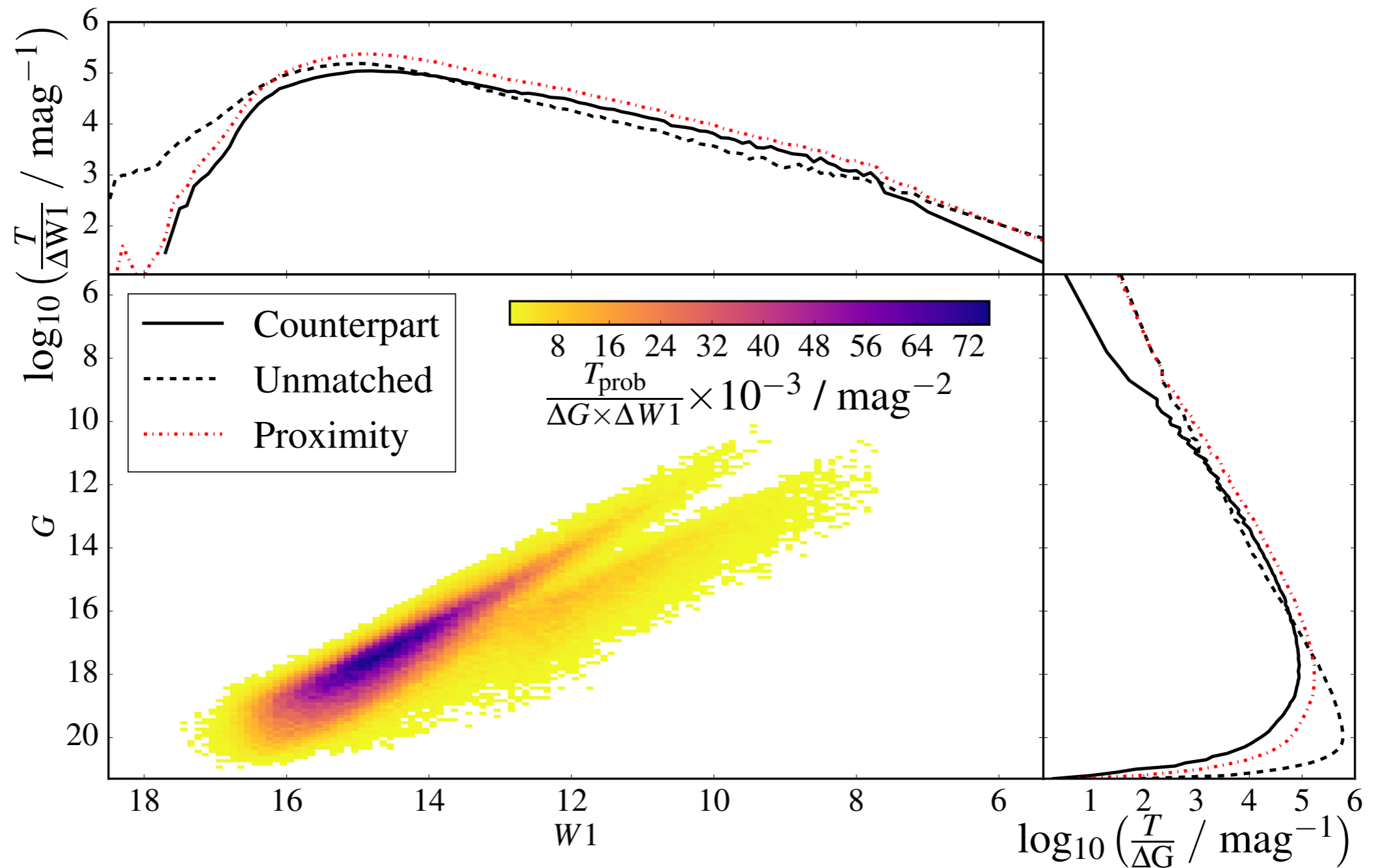
Wilson & Naylor, 2018, MNRAS, 473, 5570

Wilson & Naylor, 2018, MNRAS, 481, 2148

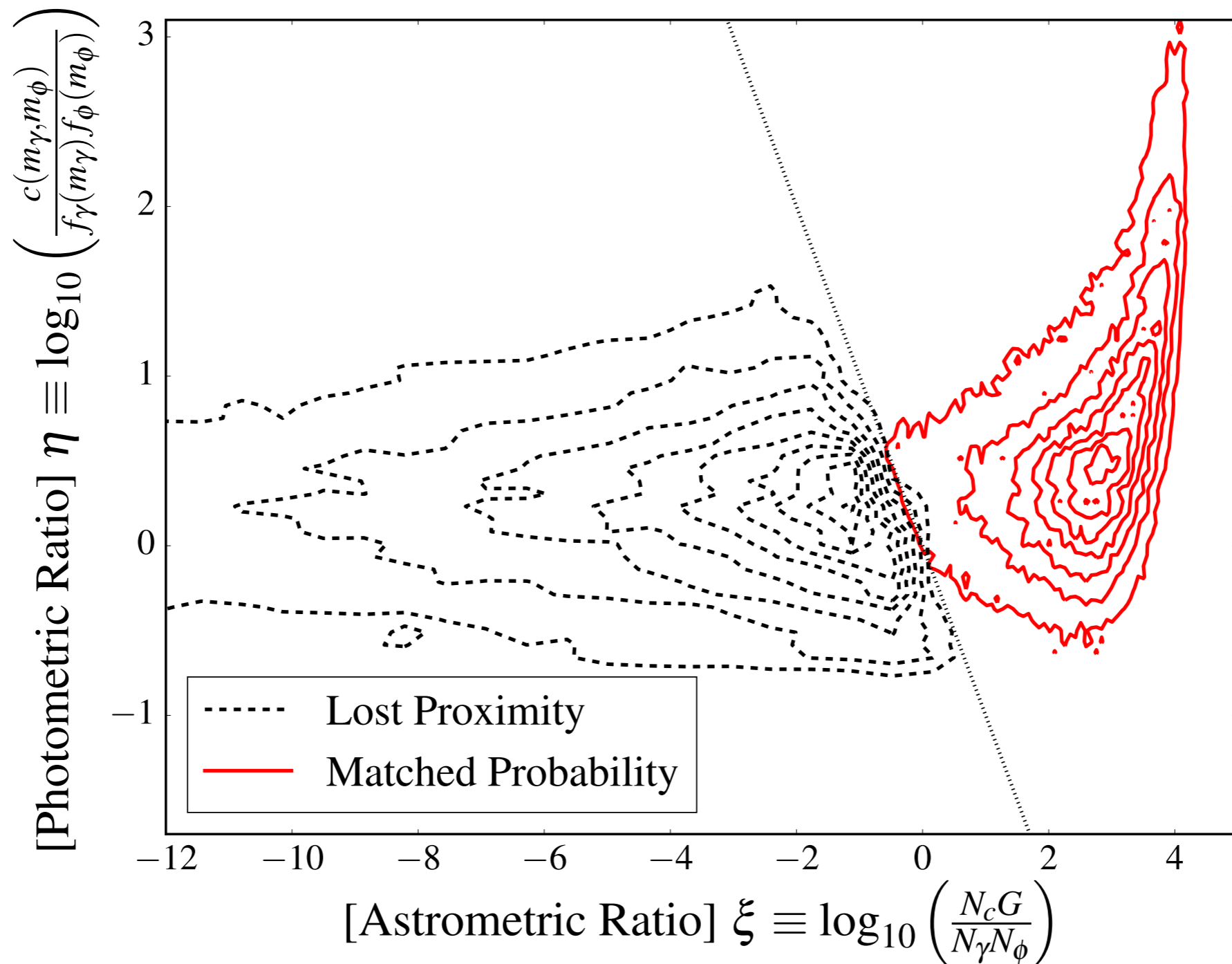
The Astrometric Uncertainty Function: Synthetic Non-Gaussian Tails



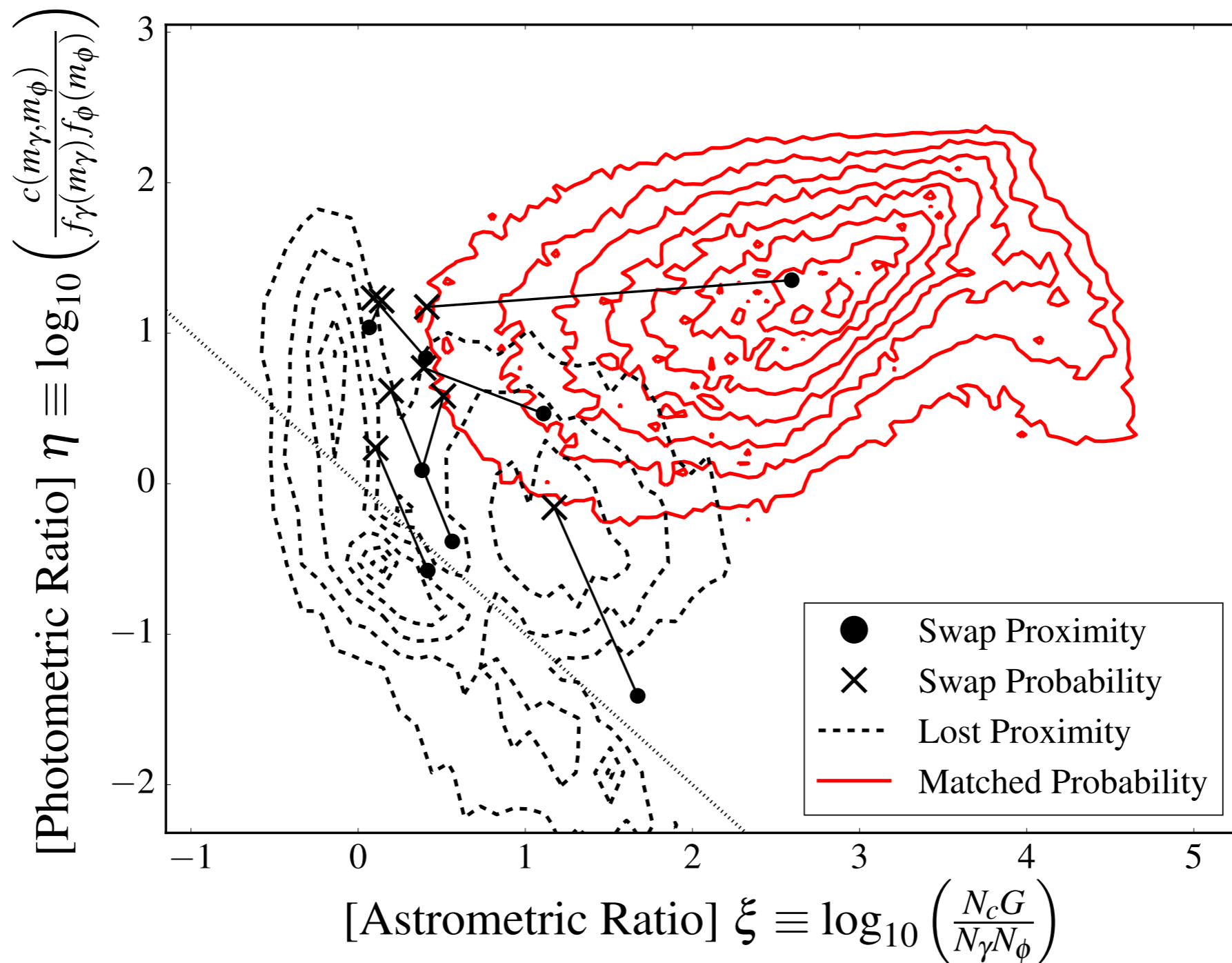
Contamination Effects: *Gaia-WISE* Gaussian Matches



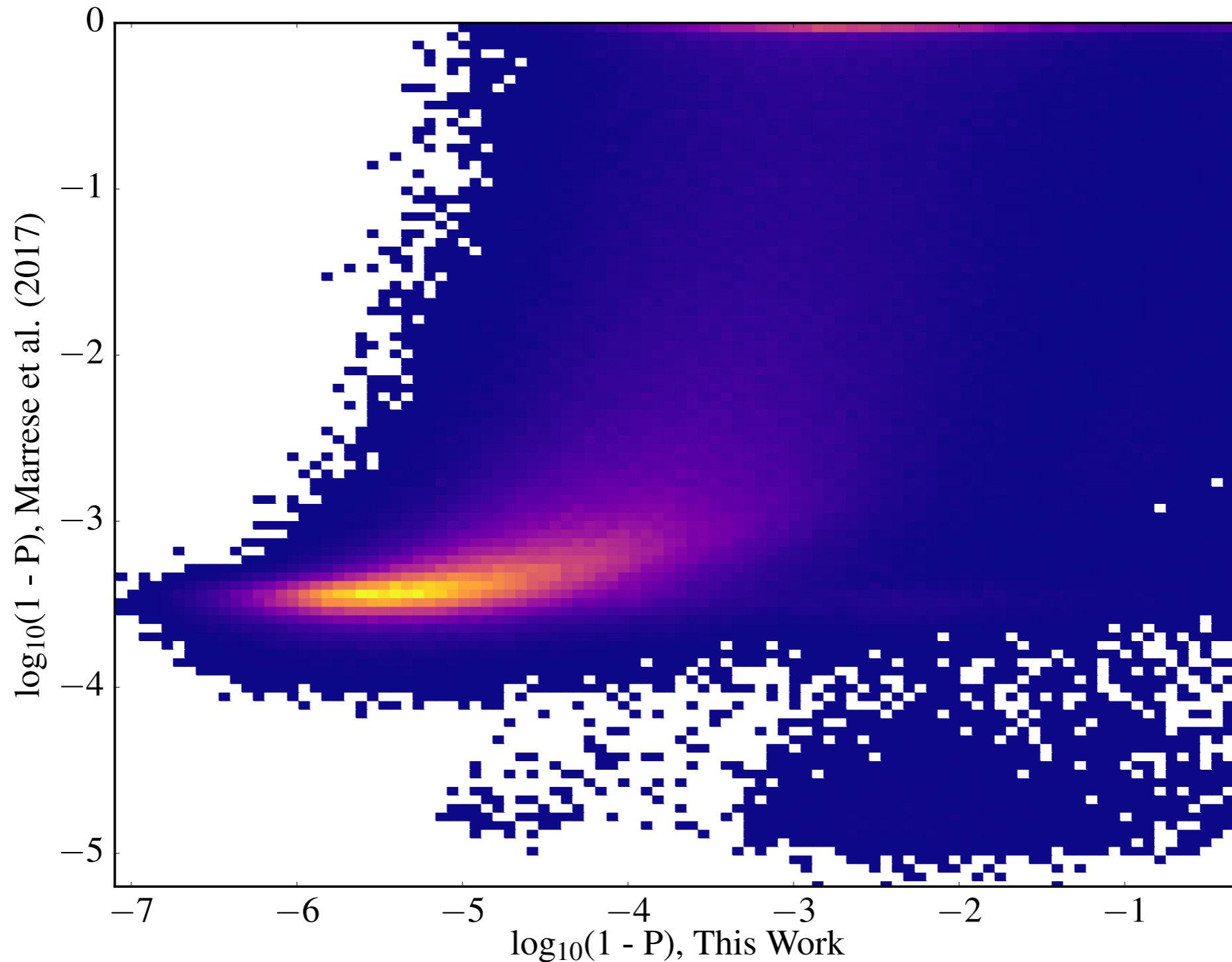
Contamination Effects: Lost Proximity Matches



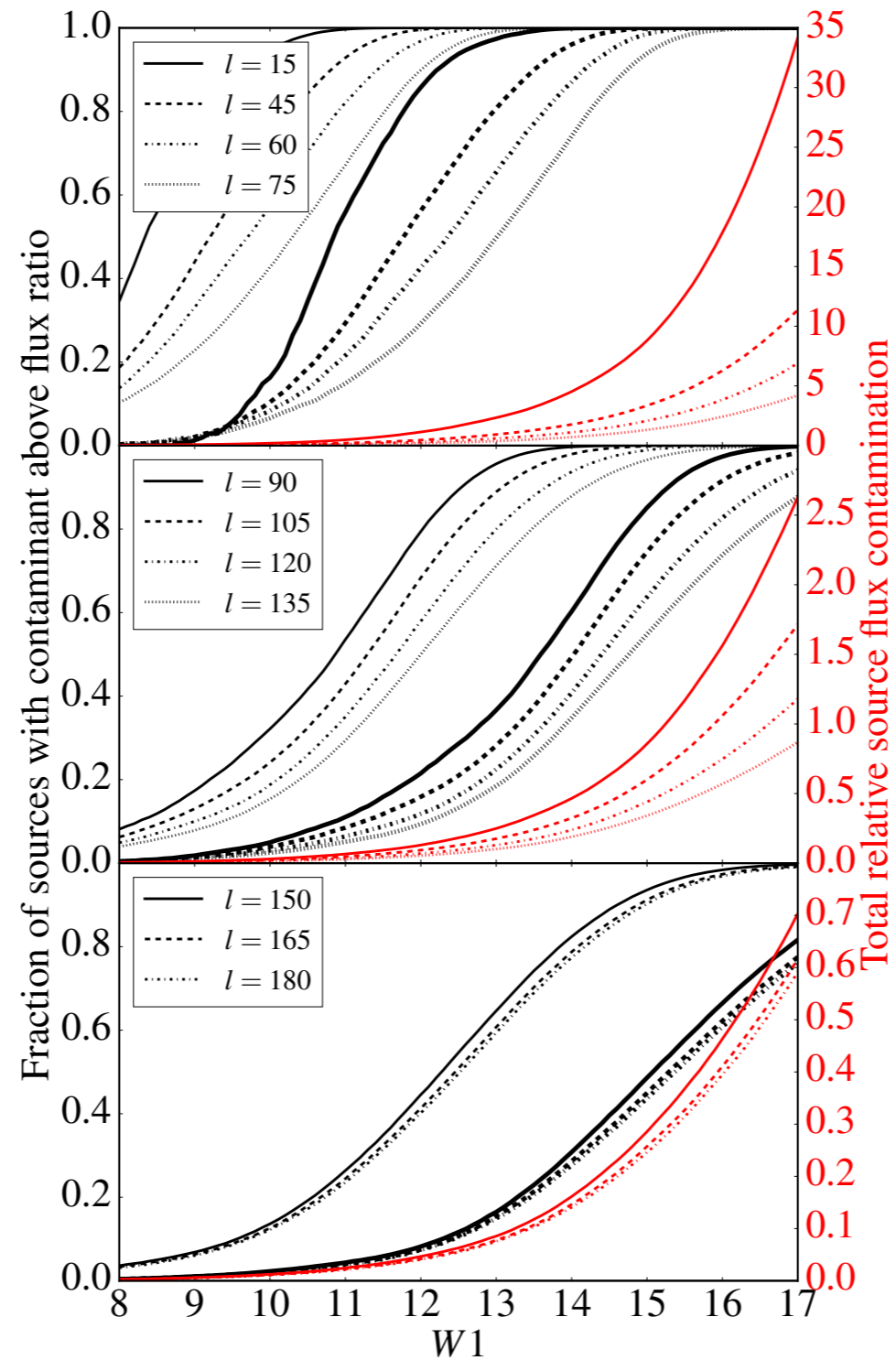
Contamination Effects: Lost Proximity Matches



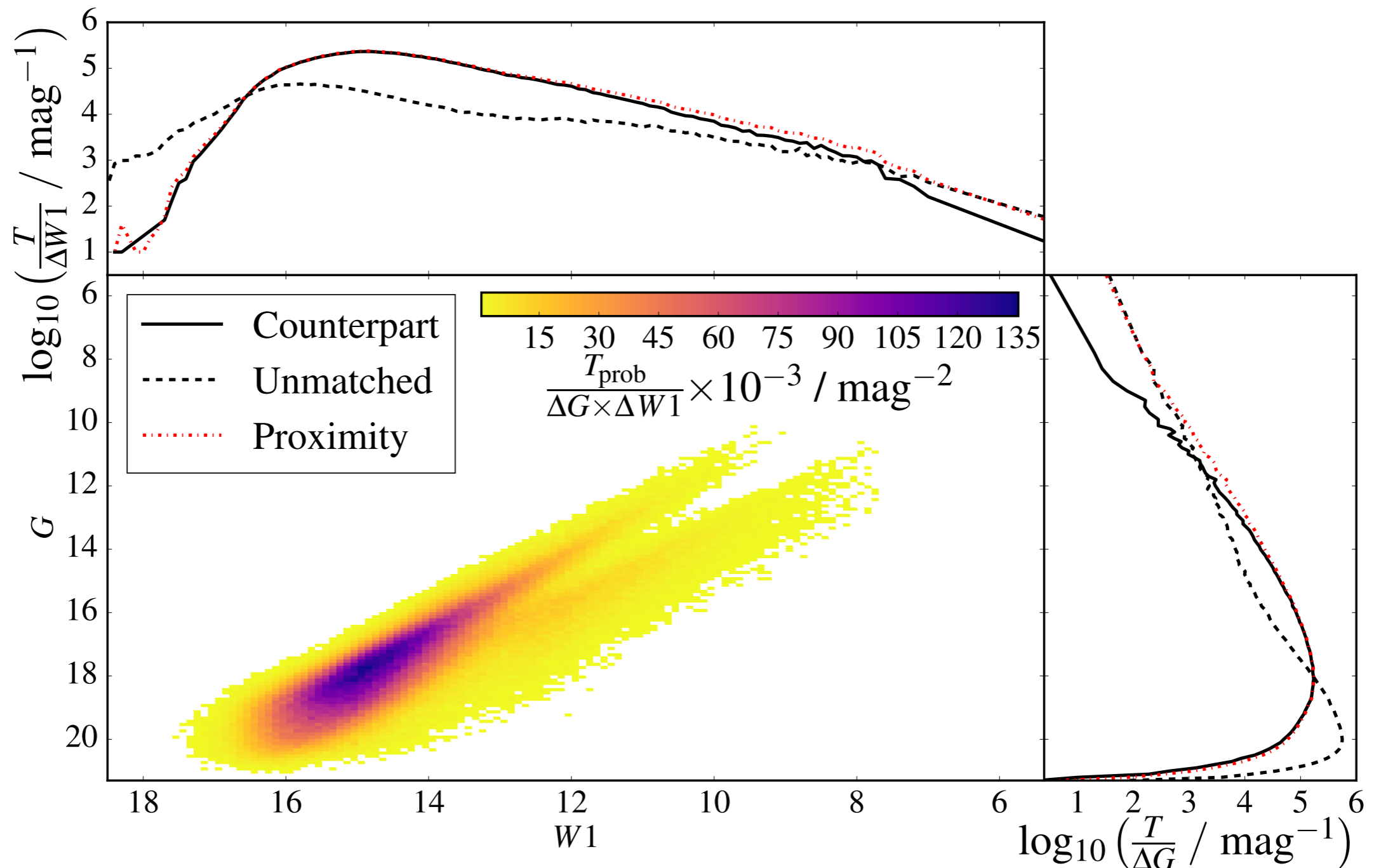
Contamination Effects: *Gaia* Lost Matches



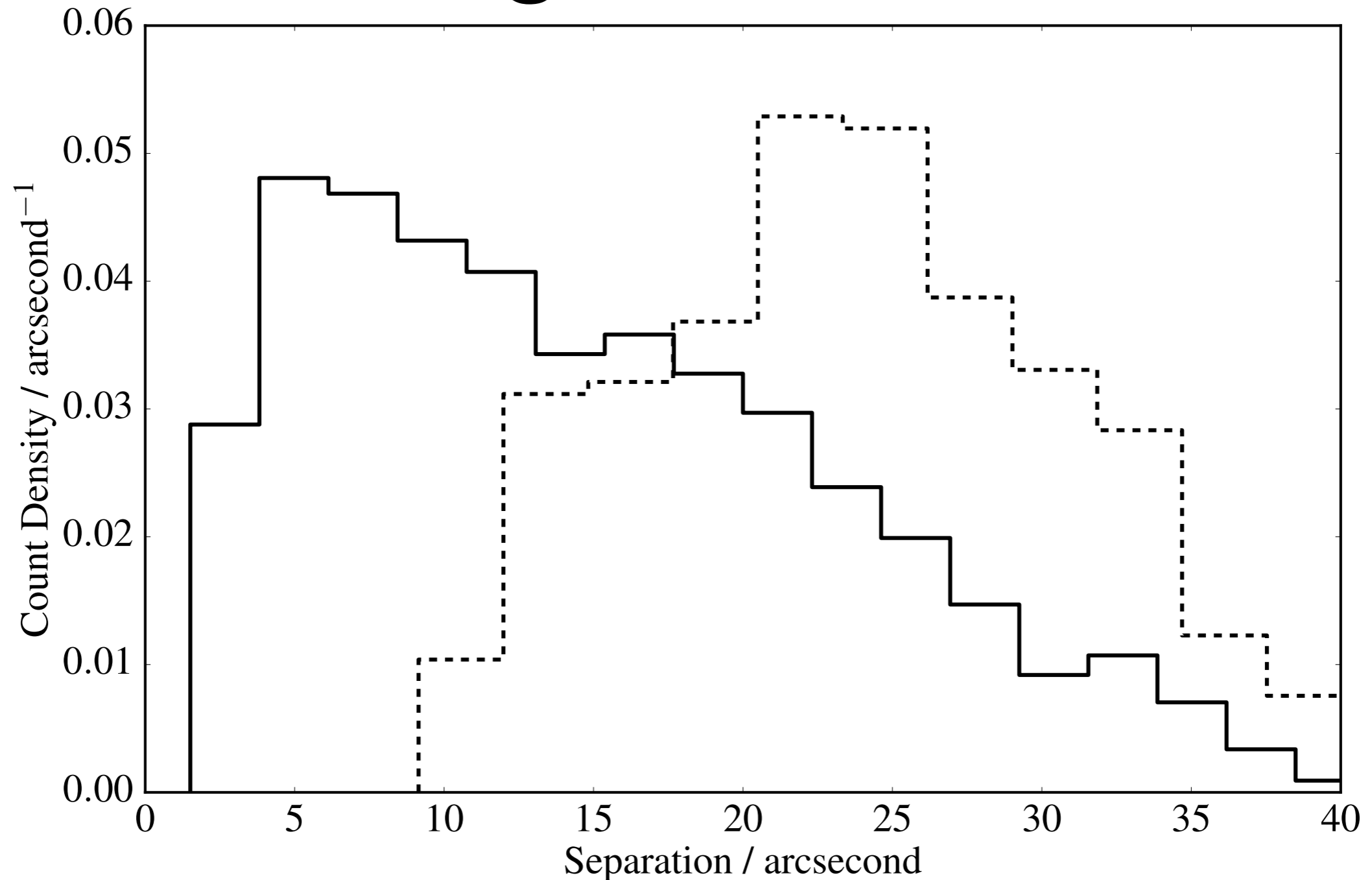
Contamination Effects: Contamination Rates & Amounts



Contamination Effects: *Gaia-WISE* Empirical Matches



Contamination Effects: Resolving Contaminants

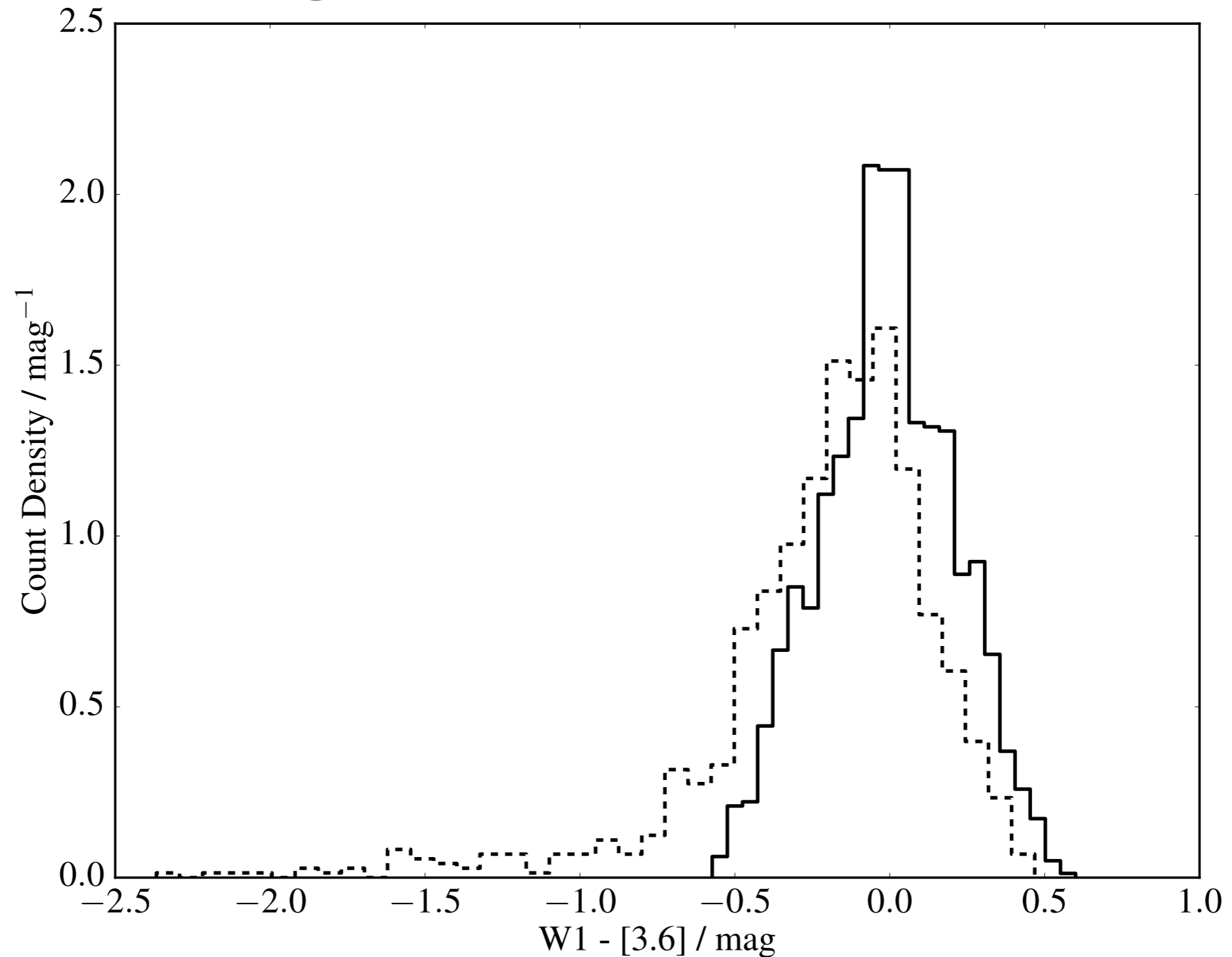


Spitzer - Werner et al., 2004 ,ApJS, 154, 1

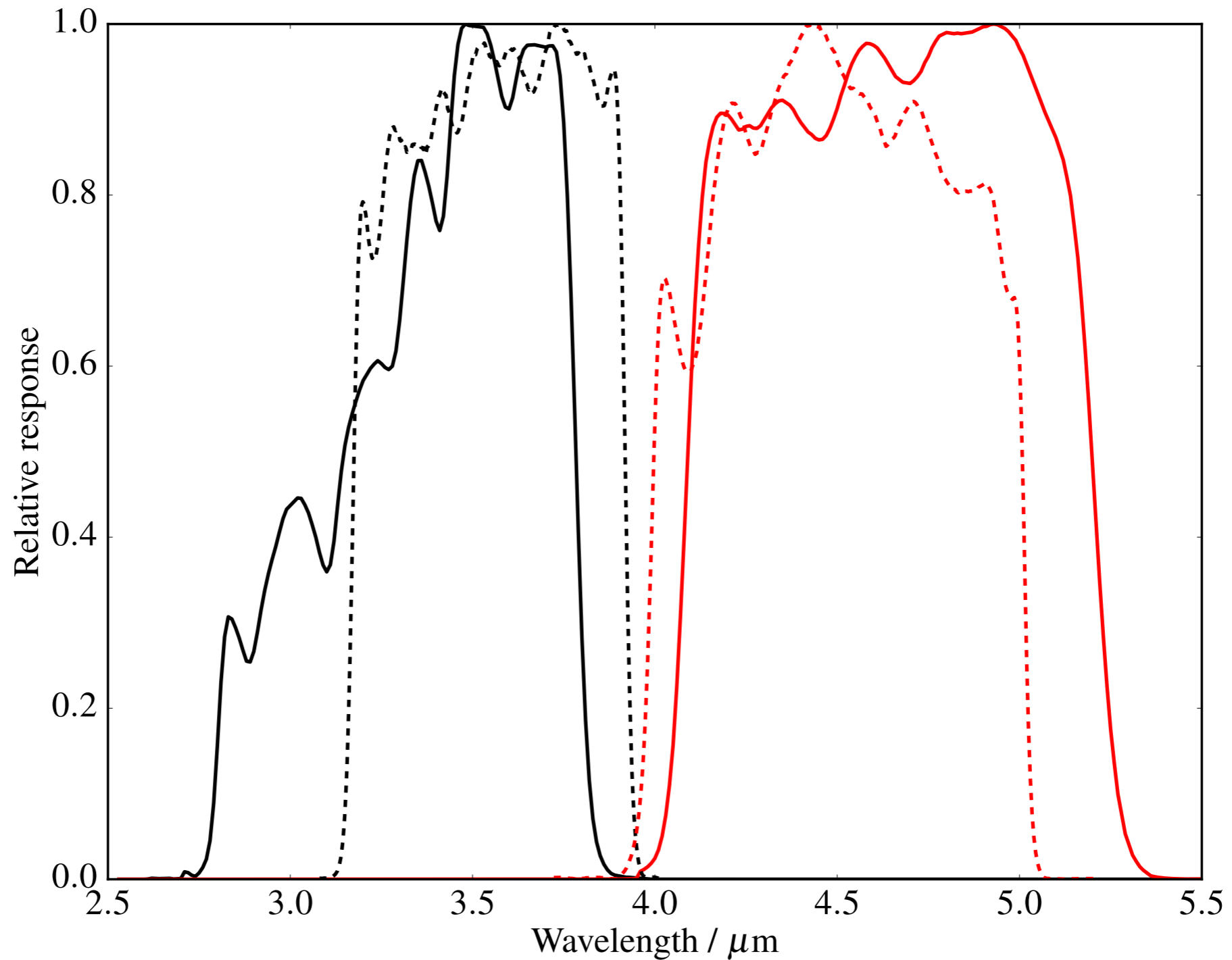
IRAC - Fazio et al., 2004, ApJS, 154, 10

Wilson & Naylor, MNRAS, 2018, 481, 2148

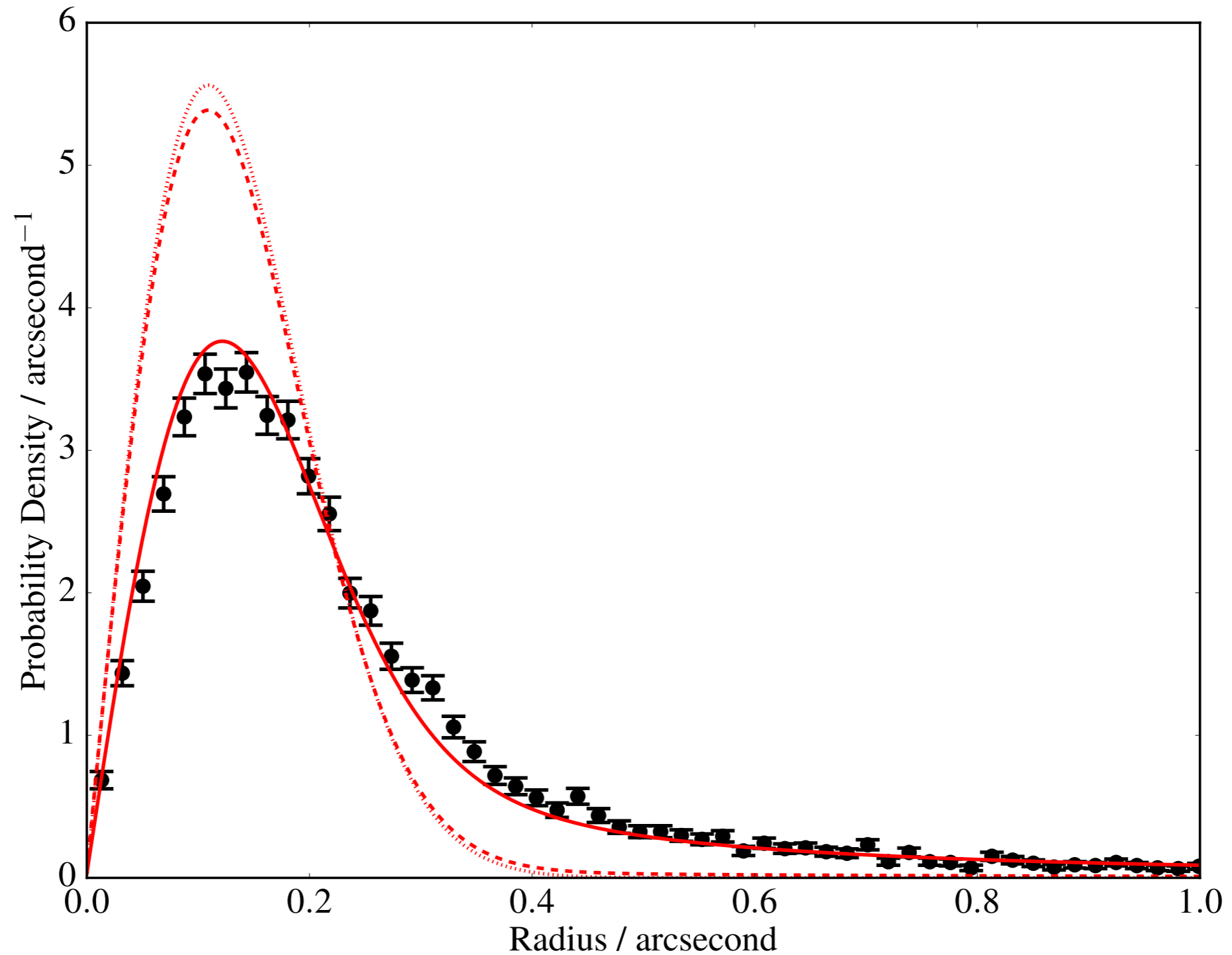
Contamination Effects: Resolving Contaminant Flux



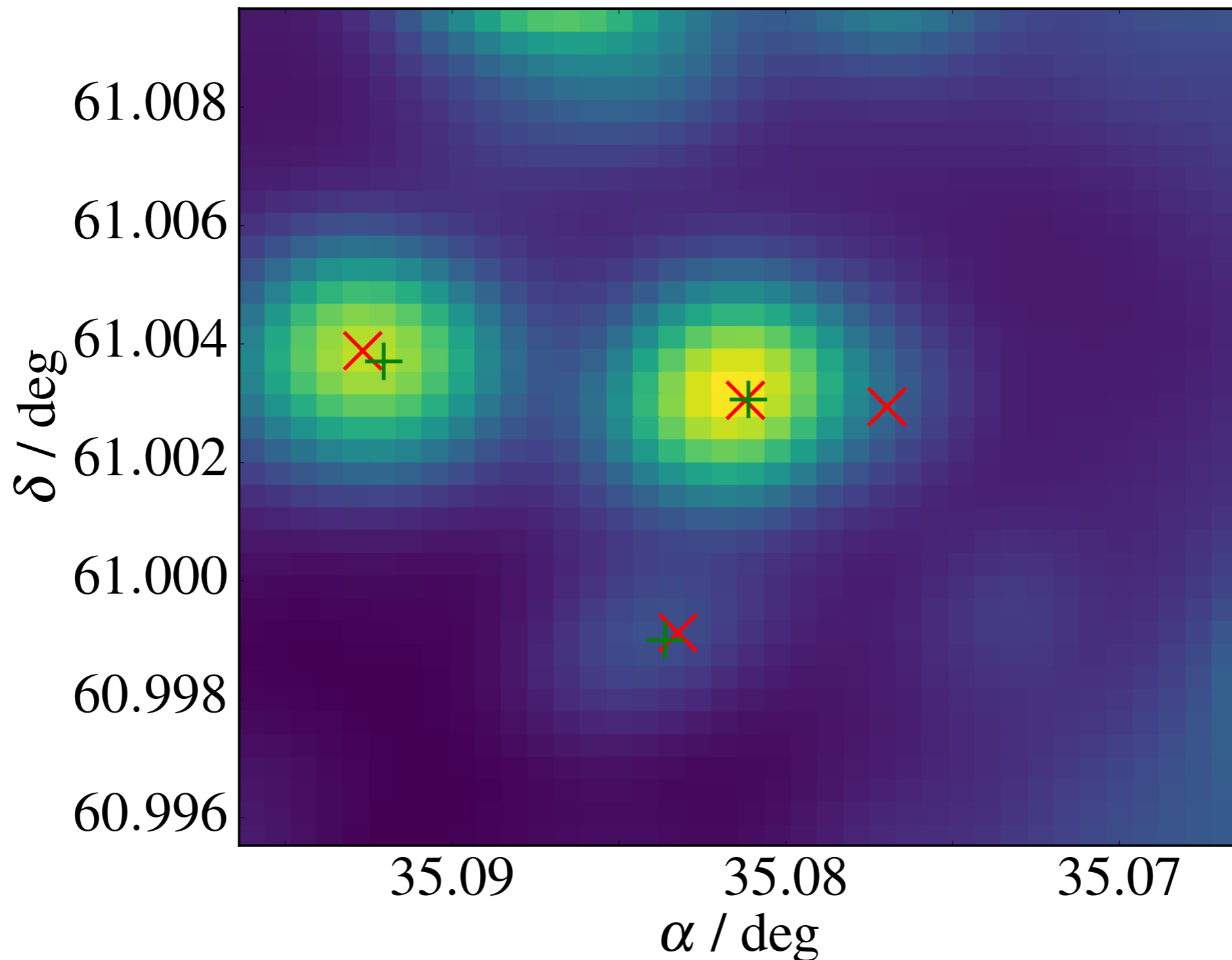
Contamination Effects: Wavelength Coverage



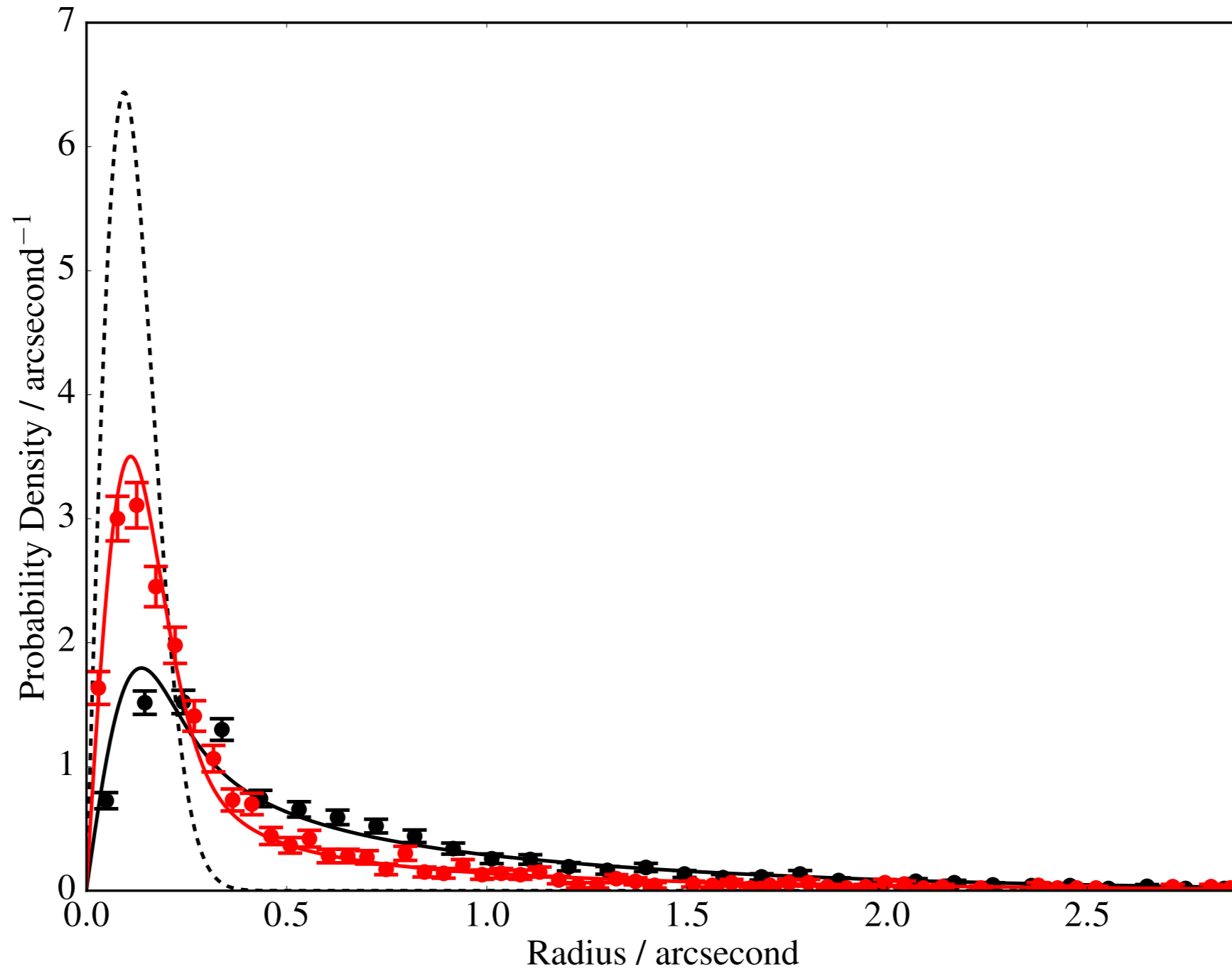
Contamination Effects: Galaxy Contamination



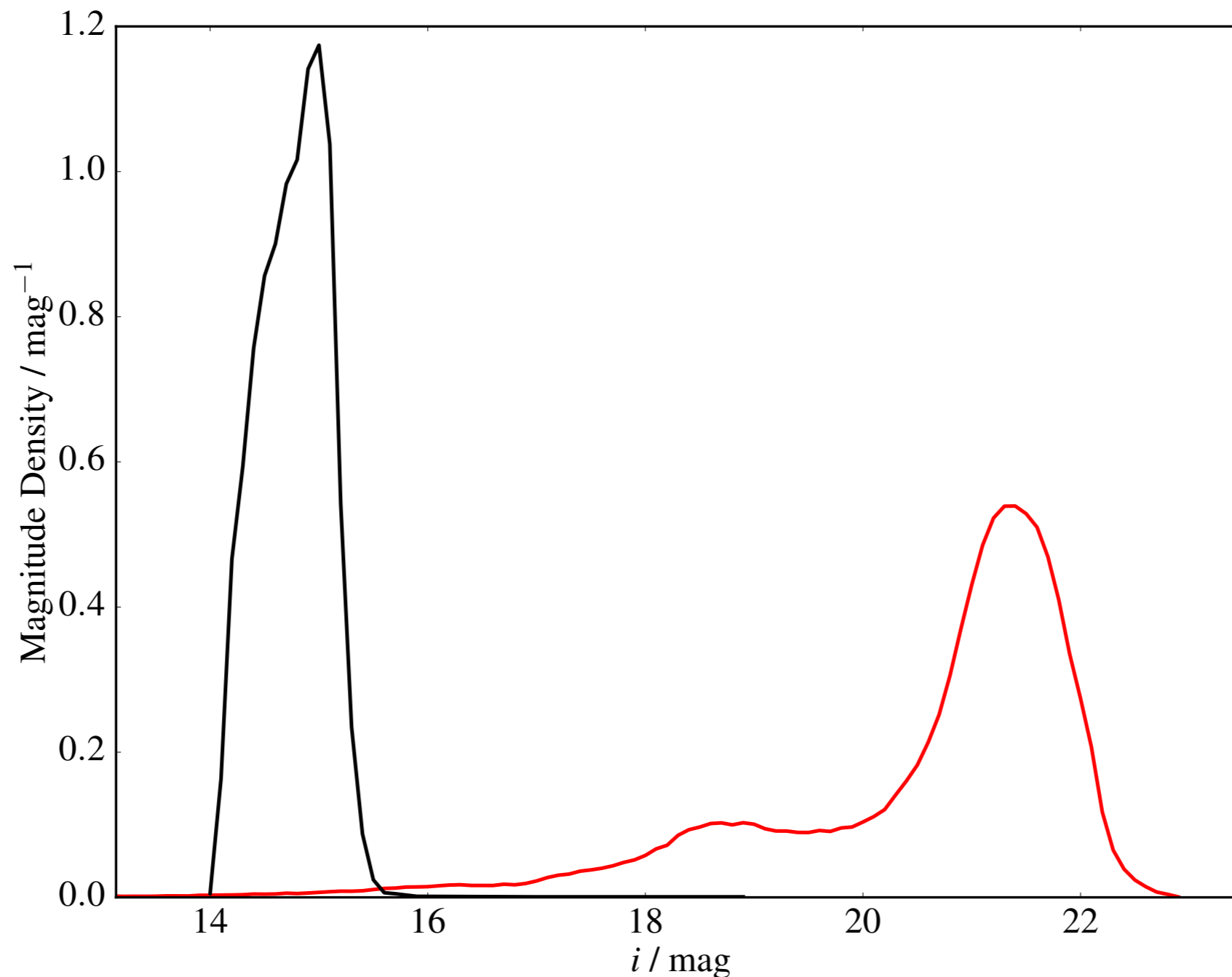
Contamination Effects: *Gaia*-Wise Resolved Blend



Contamination Effects: Crowding Normalisation



Probability-Based Catalogue Matching: Colour-based False Match Rejection



The Astrometric Uncertainty Function: Analytical perturbations

